# Nuclear Reactions at Intermediate and High

# Energies

by the collaboration of groups from Rio and São Paulo (CBPF and IFUSP)

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#### SUMMARY

- Brief description of the reaction process;
- Intranuclear cascade;
- Evaporation-fission;
- Main features of the CRISP model;
- Main reasults;
- Conclusions.

#### The main steps of the nuclear reaction

**Primary interaction** 

#### Intranuclear Cascade

#### Compound nucleus formation

Nuclear evaporation and fission

Main features of the CRISP code it inherits many of its features from the MCMC and MCEF codes

- Realistic time-ordered sequence of events during the intranuclear cascade;
- Each interaction is taken into account according their known cross section and angular distributions;
- Barion resonance formation, propagation and decay;
- Resonant and direct pion production;
- Kaon production.

M. Gonçalves et al. PLB 406 (1997) 1. S.B. Duarte et al. PRL49 (1982) 536 Deppman et al. PRL87 (2001) 182701; CPC 145 (2002) 385

#### **CRISP: main features**

- Quasi-Deuteron photoabsorption mechanism;
- Realistic Pauli blocking mechanism calculation based on the Fermi gas model;
- Photon hadronization;
- Shadowing effect;
- Vector meson production;
- One or two pions direct production in nucleonnucleon interactions during the cascade.

Deppman et al. PRC66 (2002) 067601; NIMB 211 (2003) 15;

#### Intranuclear Cascade main primary reaction

 $\gamma p \rightarrow N^* \rightarrow (p \pi^0) (n \pi^+) (N^* \pi)$   $\gamma n \rightarrow N^* \rightarrow (n \pi^0) (p \pi^-) (N^* \pi)$   $\gamma N \rightarrow \pi N$   $\gamma N \rightarrow \Lambda K$  $\gamma N \rightarrow \rho N$ 

#### **CRISP: main features**

- Particles evaporation by Weisskopf's statistical model;
- Evaporation of neutrons, protons and alphaparticles are included;
- Fission according the Bohr-Wheeler model;
- Nuclear mass calculation according droplet model;
- Separation energies are calculated from the mass formulas;

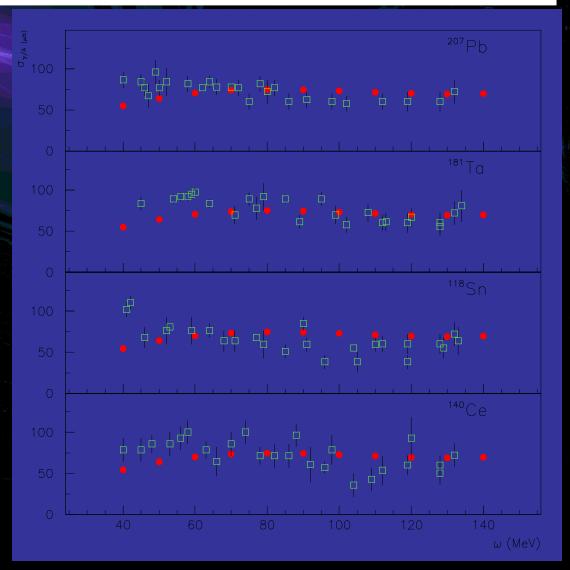
#### **CRISP: main features**

- Level density parameter after neutron evaporation is calculated by Dostrovsky parametrization;
- Level density parameters for proton and alpha evaporation following fenomenological formulas;
- Tunneling effect included in the Coulomb potential according LeCouteur;

### Main results with CRISP

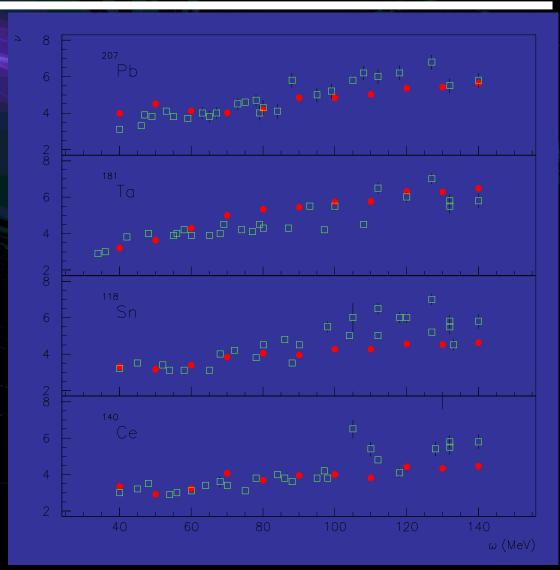
Photoabsorption cross sections at the quase-deuteron region.

> Deppman et al., Journal Phys. G: NPP30 (2004) 1991

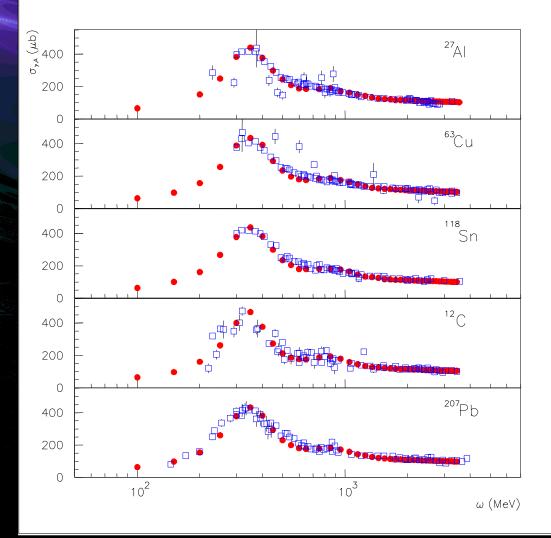


### Main results with CRISP

Neutron multiplicity after photoabsorption at the quase-deuteron region.

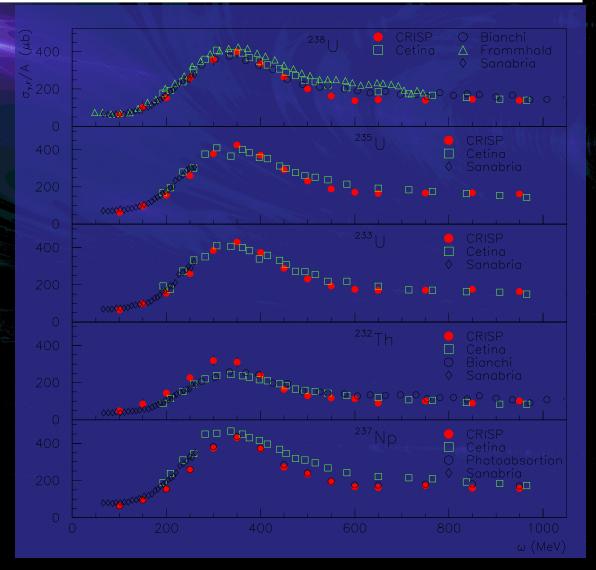


Photoabsorption cross section between 100 MeV and 3.5 GeV.

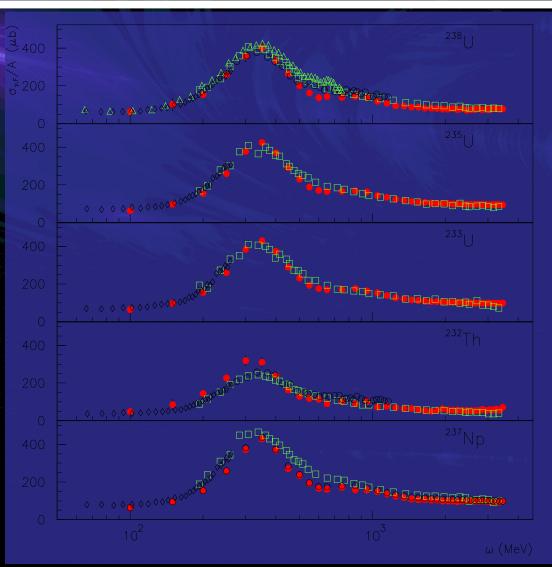


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Photofission cross section for actinide between 60 MeV and 1.0 GeV.

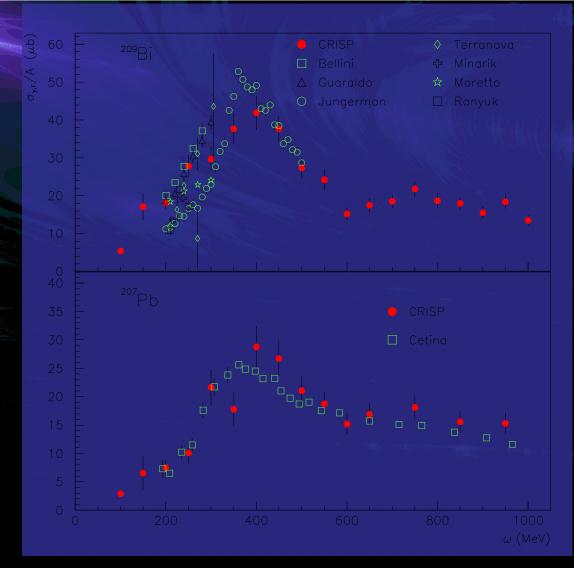


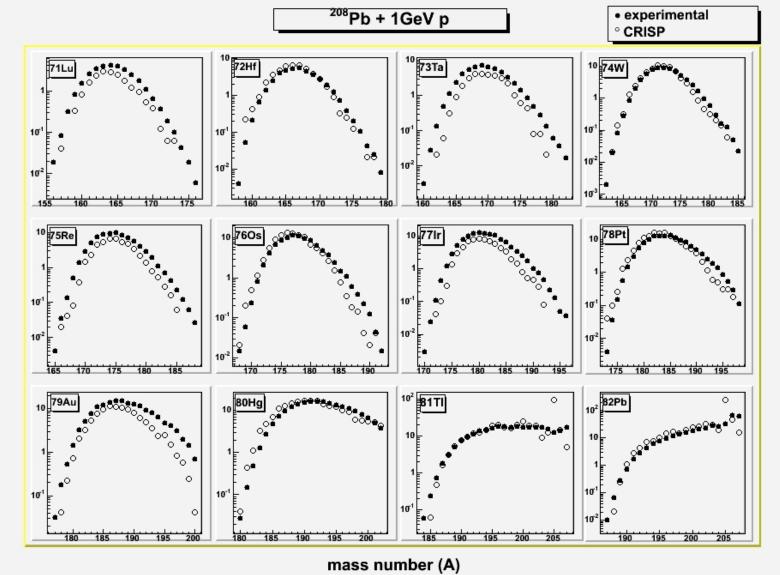
Phtofofission cross section between 60 MeV and 3.5 GeV.



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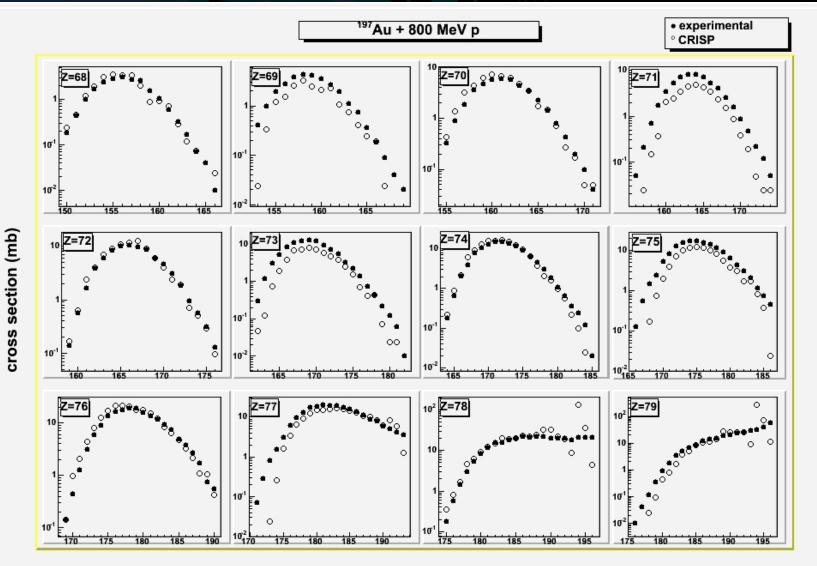
Pre actinide photofission cross section between 150 MeV and 1 GeV.





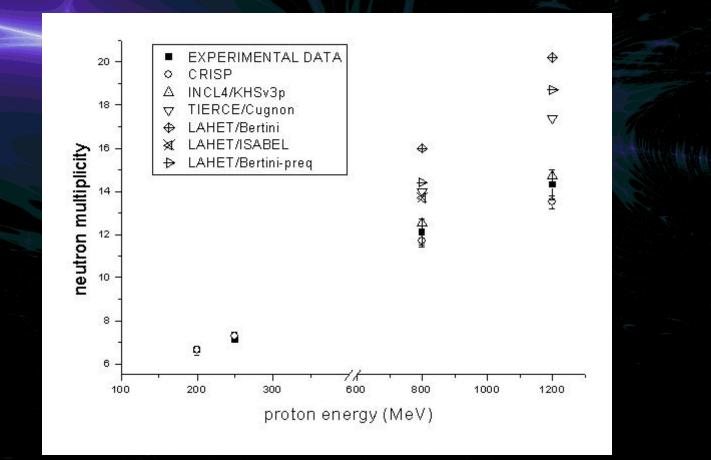
cross section (mb)

15



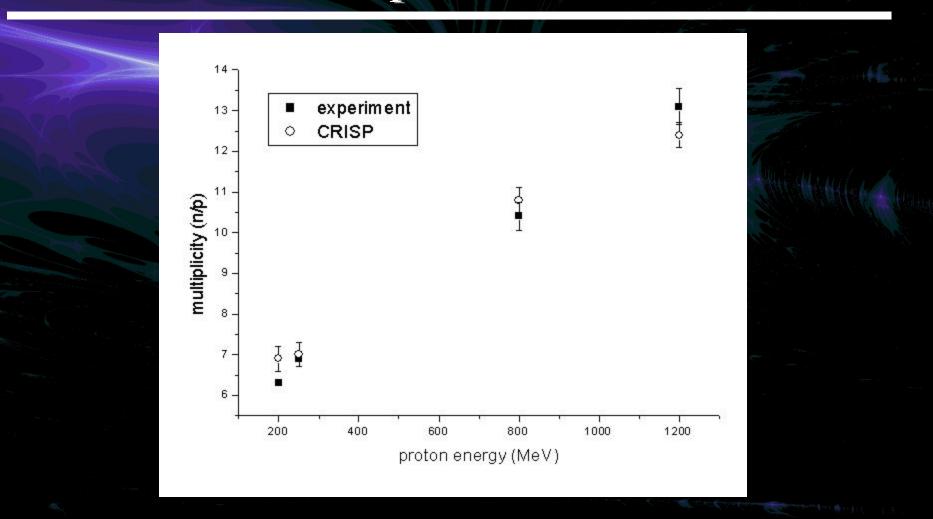
mass number (A)

#### p + Pb



Nucl. Sci. Eng. – accepted for publication.

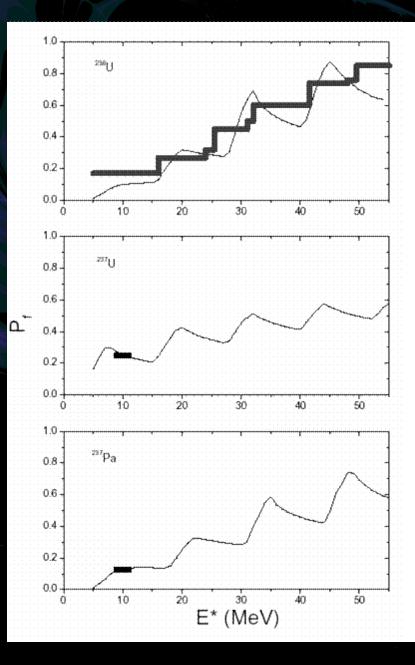
### p+Au



# Espalhamento (e,e'f)

V. Likhachev et al. PRC68 (2003) 014615

 $\Lambda$  production, propagation and decay in nuclear matter.



## Conclusions

- A Monte Carlo calculation for nuclear reactions has been developed;
- Time-ordered sequence in the intranuclear cascade;
- Works well for photon, electron and proton probes with energies from 40 MeV up to 3.5 GeV and nuclei from Al to Np;
- It has been applied for nuclear reactions studies and for the development of hybrid reactors;

## Conclusions

• It is not an event generator, but tries to simulate in a realistic way the process that take place during the nuclear reaction;

 Today we have people from São Paulo (IFUSP and IPEN), Rio de Janeiro (CBPF and IRD), Ilhéus (UESC), Havana (InsTec), Frascati (INFN-LNF) developing or applying the CRISP for simulation of nuclear reactions.