

**César Lattes e as técnicas de produção e detecção de mésons:  
a prática científica como objeto histórico**

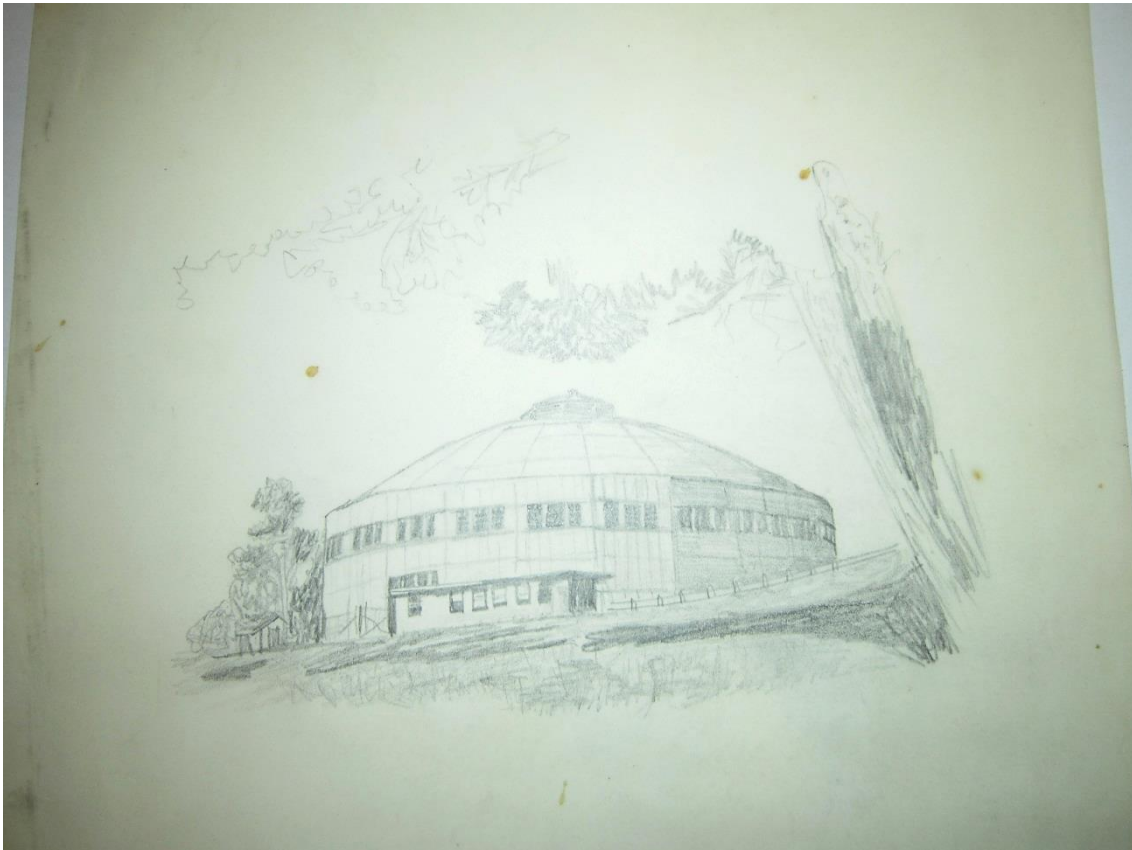
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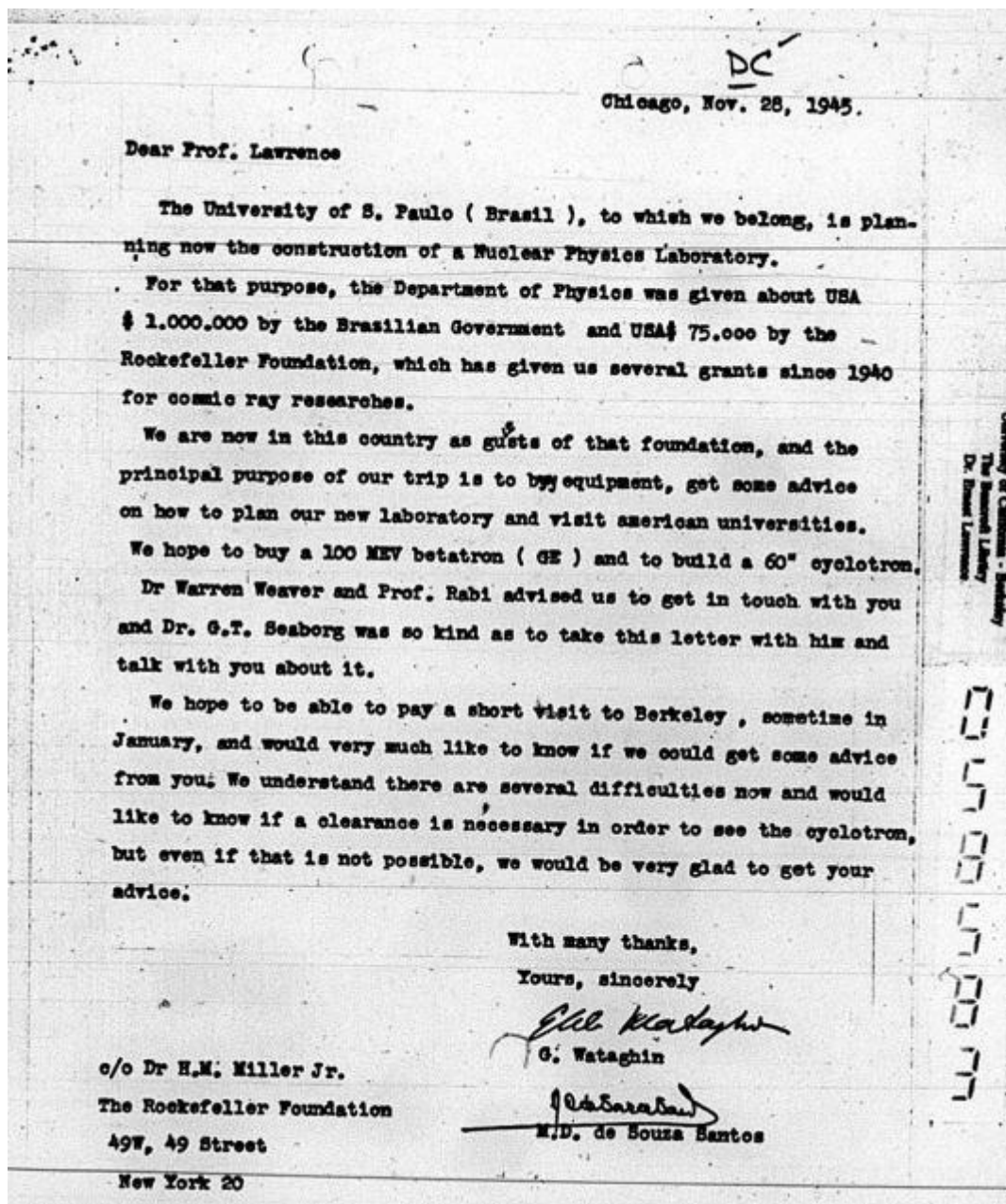
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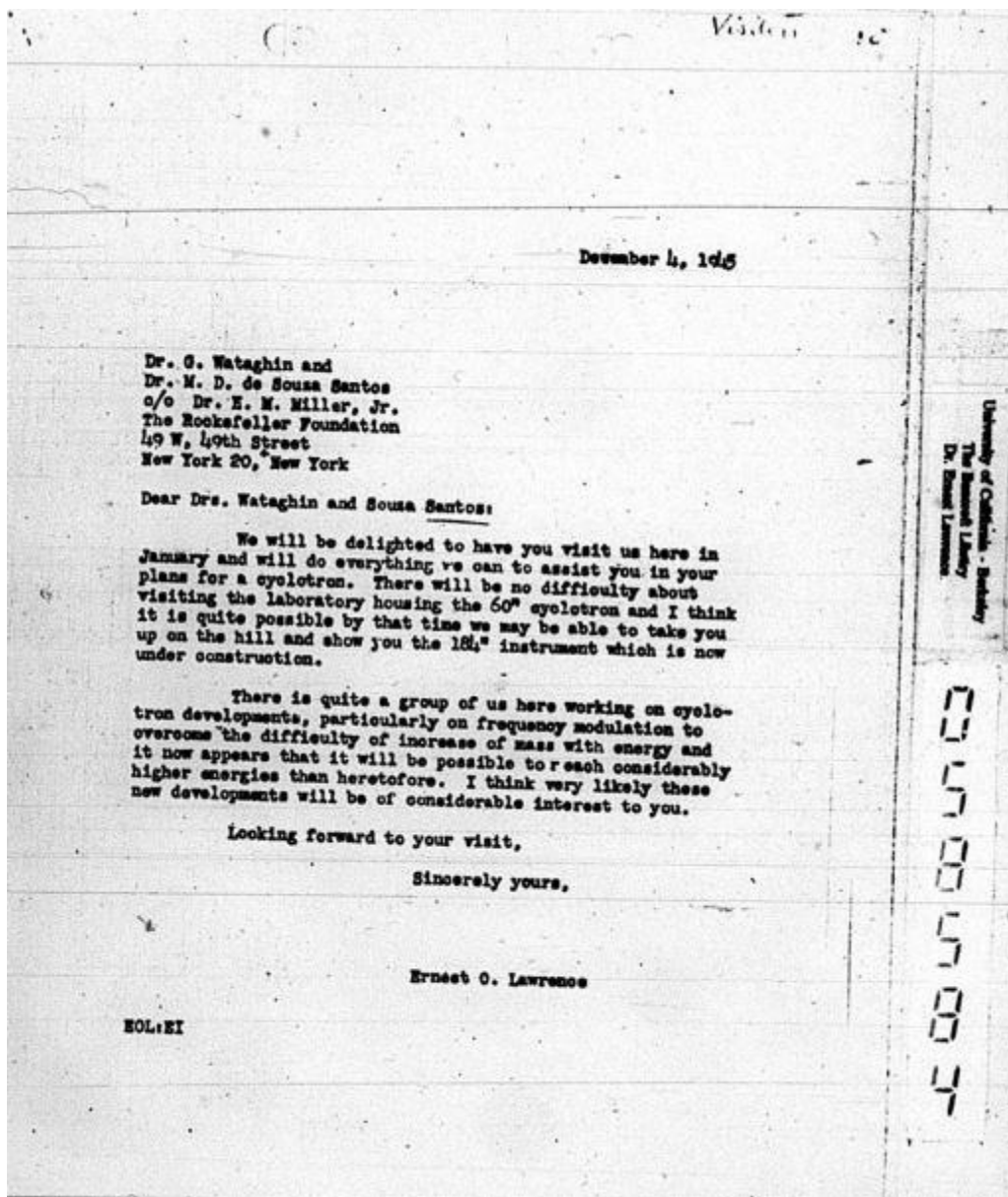
[13] E. Gardner, Lawrence Radiation Laboratory (Research and Development Records and Administrative Files of Eugene Gardner, National Archives and Records Administration, San Bruno), pasta 30, caixa 2.



[42] G. Wataghin e M.S. Santos para E. Lawrence (The Bancroft Library, University of California, Berkeley, 1945), correspondência, rolo 27, caixa 18, pasta 13.



[45] E. Lawrence para G. Wataghin e M.S. Santos (The Bancroft Library, University of California, Berkeley, 1945), correspondência, rolo 27, caixa 18, pasta 13.



[51] J.L. Lopes, em: César Lattes 70 anos: a nova física brasileira, editado por A. Marques (CBPF, Rio de Janeiro, 1994), p. 72-73.

Física Teórica e Física Superior na F.N.Fi. Eis a carta que me escreveu Lattes naquele ano:

Bristol 7- Março-1946

Caro amigo Leite,

Estou trabalhando no H.H.Wills Physical Laboratory da Universidade de Bristol. Cheguei há poucos dias e comecei a trabalhar com o Prof. C.F. Powell sobre o método das chapas fotográficas aplicado à física nuclear. Vamos iniciar experiências sobre o scattering neutron-proton, em condições muito melhores do que as dos trabalhos anteriores.

Lí seu trabalho, publicado na Revista da Academia Brasileira, sobre o deuteron, e estou muito interessado no seu trabalho sobre o scattering. Um dos teóricos que aqui trabalham está estudando o mesmo assunto. mostrei-lhe o trabalho do deuteron e, em particular a nota final, sobre o outro trabalho a ser publicado, e ele me pediu para ver se consigo arranjá-lo o mais cedo possível. Poderá você enviá-lo por via aérea ?

Meus parabéns pelas suas atividades nos E.E.U.U.. Faço votos para que você continue suas pesquisas aí no Rio, apesar da falta de ambiente

( você poderá criá-lo).

Logo que tenha resultados interessantes sobre o scattering, se for permitido, comunicar-lhe-ei. Os trabalhos anteriores não tinham evidência experimental suficiente.

Minhas recomendações ao pessoal da Faculdade e um abraço do amigo

Lattes

[51] J.L. Lopes, em: César Lattes 70 anos: a nova física brasileira, editado por A. Marques (CBPF, Rio de Janeiro, 1994), p. 73-78.

Bristol 21 de Abril de 1946  
H.H. Wills Physics Lab.  
Royalfort

Dr. José Leite Lopes  
Fundação Getulio Vargas  
Núcleo de Matemática  
Praia de Botafogo 186  
Rio de Janeiro-Brasil

Caro Leite:

Recebi somente ontem sua carta de 27/3/46. Muito obrigado pelas informações. Vejo, com prazer, que você trabalhou "no duro" em Princeton. Ainda não pude ler com o devido cuidado o apanhado geral do seu trabalho sobre o N-P scattering e forças nucleares mas vou fazê-lo logo que tenha oportunidade, pois estou muito interessado no problema.

As experiências de Champion e Powell, de fato, confirmaram os resultados de Amaldi & Cia., mas não apresentavam evidência experimental (estatística) suficiente. Essas experiências foram repetidas por Powell, Heitler e Occhialini e as medidas estão em andamento. O nº de tracks medidos até o presente momento não é suficiente para tirar-se uma conclusão segura (cerca de 900 tracks) mas toda a evidência é a favor de scattering isotrópico no sistema baricêntrico, contrariamente aos resultados anteriores. Estas novas experiências foram feitas com neutrons de 9 e 13 MeV obtidos de reações D-n provocadas pelos Deuterons do ciclotron de Liverpool. Os novos resultados merecem mais confiança do que os anteriores pois a experiência foi feita com uma técnica melhor do que no caso anterior e estão sendo

∴ A evidência  
estatística  
na  
font. 13/46

levadas em conta todas as correções necessárias em vista da geometria da experiência. Amaldi está informado do progresso das experiências feitas aqui e creio que o grupo de Roma vai repetir as experiências que você conhece (naturwiss.). Ferretti esteve aqui para discutir os resultados mas isso foi há alguns meses, isto é, antes de termos a atual evidência. Havia um ponto fraco nos cálculos das experiências do grupo Amaldi mas Ferretti, que é o responsável por toda a "calculeira", após três dias de verificações febrís, chegou à conclusão de que os resultados não são alterados e que está tudo O.K. Como você vê, a situação está bastante confusa e meu conselho é que será melhor, para você, ignorar os resultados experimentais até mais completa evidência. Logo que haja mais certeza avisá-lo-ei.

Eu não estou trabalhando no N-P scattering, por enquanto, pois trata-se de experiências já iniciadas há muito tempo. Logo que o Ciclotron de Liverpool esteja em condições, iniciaremos (Powell, Heitter, Occhialini e eu) o estudo do scattering de neutrons de 17 MeV da reação B+D (de 5MeV do Ciclotron). Nessa experiência usaremos um novo tipo de placa que permite uma precisão muito melhor do que com as antigas. Vou mandar-lhe separata de um artigo que vai sair dentro de poucos dias no Journal of Sc. Instr., descrevendo as "virtudes" das mesmas e que, segundo posso garantir, vai eliminar o scepticism que ainda existe em alguns meios, sobre a aplicação das chapas fotográficas aos problemas da física nuclear.

Somente um físico daqui, Ramsey, está trabalhando na parte teórica do scattering N-P. Ele está usando a teoria de Møller-Rosenfeld. Parece-me, todavia, que ele é "rather obtuse".

Meu trabalho aqui está dividido em várias partes e tem sido, por enquanto, quase que puramente experimental. Em Bristol estou estudando e desenvolvendo a técnica da medida automática nas emulsões. Iniciei, também, uma experiência para a determinação da radioatividade do Samario; trata-se de uma terra rara que emite espontaneamente partículas alfa e, segundo uma fraca evidência experimental, protons. Com as placas que tenho à disposição o problema será resolvido facilmente. Em Cambridge estou estudando reações provocadas por deuterons de 1 MeV produzidos pelo gerador de alta

tensão. No momento estou interessado em "targets" leves (D, Li, Be, B, F); vou procurar esclarecer alguns pontos obscuros relativos a essas reações, ao mesmo tempo, determinar a relação energy-range para partículas alpha, protons e deuterons nas placas que serão usadas no futuro (até pouco tempo atrás, as placas foram tratadas com desprezo pela maioria dos físicos nucleares - veja relatório de Smith - agora a situação está mudando e está se preparando uma verdadeira corrida). Vou estudar também, desintegrações produzidas por neutron na própria placa, carregando a mesma com sais de D, Li, Be, etc. Essas experiências de Cambridge serão feitas por Cuen (um físico francês que aqui chegou pouco depois de mim, para aprender a técnica das chapas) e por mim.

Em Liverpool vou tomar parte nas experiências de scattering (N-P, P-P, D-P e scattering por gases). Essas experiências serão dirigidas pelo Powell e, provavelmente, contarão com a colaboração de Chadwick. Na parte teórica estou estudando a ação das partículas nucleares nas emulsões fotográficas. Esse estudo estou fazendo em colaboração com um físico do research lab da Kodak, que está aqui pesquisando sob a orientação do Prof. Mott. Trata-se de um jovem de grande capacidade e tenho aprendido muito com ele. Estamos estudando, também, a possibilidade de fabricar placas sensíveis a eletrons. Se isto for conseguido, será um resultado muito útil para as pesquisas sobre raios beta. Caso haja essa possibilidade, é provável que faça um estágio na fábrica Kodak em Londres, pois estou em ótimas relações com o pessoal do research lab de lá.

Como você pode ver, trabalho não falta. Mal consigo dar conta do recado. As condições de trabalho aqui são ótimas. Estou aqui a convite da Universidade de Bristol (o Occhialini conseguiu isso), recebo um ordenado mensal e tenho ampla liberdade de trabalho e iniciativa. Posso trabalhar no que mais me interessa e ficar o tempo que quiser. É uma verdadeira "mamata".

Quanto à experiência que você sugere, relativa ao inelastic scattering N-P, seria muito interessante fazê-la mas, infelizmente, não temos meios. Os neutrons mais enérgicos de que dispomos são de 17 MeV. Creio que para essa energia a porcentagem de scattering inelástico será

Estados  
líquidos

Colisões na  
curva -  
N → nucle

Scattering.

Condições  
de  
trabalho



desprezível. Seria mesmo muito difícil detectar os 3% que você calculou pois teremos background de neutrons provenientes do tank do Ciclotron, que tem um espectro contínuo. Qualquer outra sugestão que você tenha a fazer será recebida com todo o interesse por nós.

Quanto ao artigo sobre o método das chapas fotográficas, já comecei a escrevê-lo. Falei com Powell e Occhialini e é provável que seja feito em colaboração por nós três. Ficarão pronto em pouco tempo, pois temos todos os dados e fotografias. Pretendo (ou pretendemos) escrever uma coisa do tipo Rev. Modern Phys. e, se não houver nada em contrário será escrito em inglês, pois muita gente estará interessada no mesmo como fonte de dados e informações. A publicação ou não desse artigo depende ainda de circunstâncias especiais mas espero que tudo se resolva dentro desta semana. As "Summas" da fundação pedem sugestões à respeito de "improvements" eventuais que possam ser feitos nas mesmas. Sugiro, e será importante para o artigo das chapas fotográficas, que se procure melhorar a reprodução fotográfica (veja boletim da Biologia, artigo do Karen). Seria interessante se as fotografias e fotomicrografias fossem reproduzidas em papel especial (como é feito em muitas revistas científicas). Isso é particularmente importante se queremos fazer propaganda da revista de maneira a poder publicar os artigos na mesma e ter certeza de que serão lidos no estrangeiro, pois, será necessário, pelo menos no início, contar com a colaboração e interesse de alguns cientistas estrangeiros e estes não estarão muito inclinados a publicar numa revista que não tenha boa reprodução fotográfica. Outra sugestão que tenho a fazer, e que com certeza deve ter ocorrido a vocês, é que, uma vez que a Summa da fundação é a única revista de física no Brasil, seja criada na mesma a seção de cartas ao Editor ou notas prévias, cuja utilidade é bem conhecida. Já falei com o Powell e Occhialini sobre o convite para colaboradores permanentes. Eles, assim como o abaixo assinado, sentem-se honrados, aceitam e agradecem. Vou convidar, também, Roseblum (físico francês que trabalhou no espectro de estrutura fina das partículas alfa), e outros, quando aparecer a oportunidade. Saiba me dizer se a fundação concorda.

A separata do Dirac seguirá apenas a consiga. O Mário tem uma e eu emprestei (ou dei, não me lembro) a minha ao Freire. Vou, também, lhe enviar separatas dos trabalhos de Born e Peng. Quanto aos de Powell enviar-lhe-ei um microfilme, pois estão esgotadas. Livros bons são coisa rara aqui; imagine que não consigo encontrar o Dirac, Heitler, Gamov etc. A guerra aqui deixou uma miséria miserável. ✦

Soube com prazer que você está na Fundação. Pelo que tive oportunidade de observar o pessoal de lá está com uma ótima orientação e animados a fazer alguma coisa séria pela Ciência no Brasil. Queira transmitir minhas recomendações aos Drs. Assiz Ribeiro, Othon Leonardos, Manoel Ferreira e Charles Javes, assim como ao pessoal do núcleo de matemática.

Ficar-lhe-ia grato se você mandasse cópia da parte da carta em que dou informações sobre as experiências do scattering N-P ao Mario, pois sei que ele está interessado no problema e prometí escrever. Pergunte a ele quais são as separatas de que mais precisa, que procurarei manda-las. Vou escrever a ele por estes dias.

A impressão que tivemos, o Occhialini e eu, é de que voce vai aguentar firme no trabalho e não se deixará influenciar pelo "clima tropical". São esses, também, os votos que fazemos. De minha parte, espero poder colaborar quando voltar, ou mesmo aqui: às ordens.

Tive uma idéia. Talvez seja muita fantasia, mas, em todo caso, lá vai. O que pensa voce de escrevermos um livro em português sobre física atomica e nuclear: dois volumes, talvez. Não há nada sobre o assunto no Brazil e parece-me que seria interessante começar. Voce poderia se encarregar da parte mais teórica e eu da parte mais experimental. Não seria uma perda de tempo, pois, escrevendo, aprenderíamos muito e teríamos muitas sugestões para pesquisas eventuais. Que tal? aguardo sua resposta assim como sua opinião sobre as possibilidades praticas da ideia. Naturalmente será uma coisa difícil devido ao escasso número de leitores mas não devemos esquecer Física Nuclear é o assunto do dia. ✦

Inicialmente poderia se publicar sob forma de apostila. Como vai o Tiomno em S. Paulo? Ele vai pedir bolsa para este ano? Caso ele esteja interessado em vir para a Inglaterra, e caso ele consiga uma bolsa, diga-lhe para me escrever, pois estou em condição de dar informações sobre os

J. Leite Lopes

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melhores lugares para trabalhar aqui. Existe em S. Paulo um rapaz que está cursando o 3º ano. Hugo Camerini trabalhou comigo durante o ano passado e é um rapaz de talento. Peço-lhe que o ajude caso o mesmo precise de orientação. Parece que o Mario vai aos E.U. e o laboratório de São Paulo ficaria como você sabe.

Espero para breve a notícia fatal de seu casório. Não se esqueça de comunicar. Quanto à mim, você já deve saber das novidades... sentimentais. Provavelmente você já tinha previsto.

Bom. Por hoje é só.

Um abraço de seu amigo

Cesar Lattes

P.S. Para conseguir colaboração estrangeira é indispensável melhorar a reprodução fotográfica. A promessa de rápida publicação dos trabalhos deverá ser mantida rigorosamente.

[51] J.L. Lopes, em: César Lattes 70 anos: a nova física brasileira, editado por A. Marques (CBPF, Rio de Janeiro, 1994), p. 89-92.

Cesar Lattes, o CBPF e a Nova Física no Brasil

Bristol 15/11/47

Caro Lúcio.

Requero rapidamente, sua carta que estou apreciando. - Enviarei novamente para a semana.

Parabéns pela atividade e resultados. Trabalho muito animado.

Haverá: enviarei as fotos e via vi. do seu País dentro de 2 semanas.

Fótilos: enviarei em seguida.

n-n scattering. Pensei que seria possível formoselo uma mistura de neutrinos lentos da Pilla com hélio líquido e usando isto como target...

Duplos meus:

7 casos já encontrados, em que o mesmo sinal está terminando na emulsão. Alcanze (Range) homogêneo =  $60 \pm 8$   $\mu$ verois. Energia  $\sim 4$  Mev.

Massa do secundário definitivamente menor que a do primário (grain counting mostra isso de maneira definida). Razão de massas: ainda não temos suficiente n. de medições para dar valor preciso mas fica entre  $\frac{m_p}{m_s} = 1.3$  e  $2.2$ .

Se  $\frac{m_p}{m_s} = 1.3$  o processo é  $p \rightarrow n_s + \pi^+$  (-50-90) ou neutrino

Se  $\frac{m_p}{m_s} = 2.1$   $p \rightarrow n_s + \pi^+$  neutrino de mesma massa que  $n_s$ .

As massas absolutas também não foram determinadas com precisão mas são da ordem de  $m_s \sim 320$  e.m.

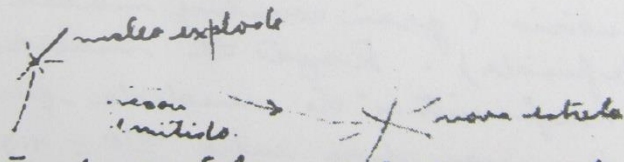
Em 400 pessoas que luminam nos emulsores  
 7 têm secundárias maior que  $500\mu$  (espessura  
 da emulsão  $5\mu \times 100\mu$ ) consideração de  
 ângulo sólido mostram que esse 7 correspondem  
 na realidade a  $\approx 120$  duplos mesons (isto é  
 o processo de  $\mu$ -decay).

65 mesons são capturados por nucléos e os  
 desintegrados, de maneira que dos 400 temos

120	primários ( $\pi$ )	} $\mu$ -meson
120	secundários ( $\mu$ )	
65	negativos capturados	

O  $\mu$ -decay é, por, muito comum e, pela  
 existência que temos, todos os mesons que não  
 são capturados sofrem  $\mu$ -decay antes de darem  
 $\beta$ -decay.

Temos 10 mesons emitidos por nucléos em  
 explosão, dois desses mesons são por sua vez  
 capturados e dão nova estrela:

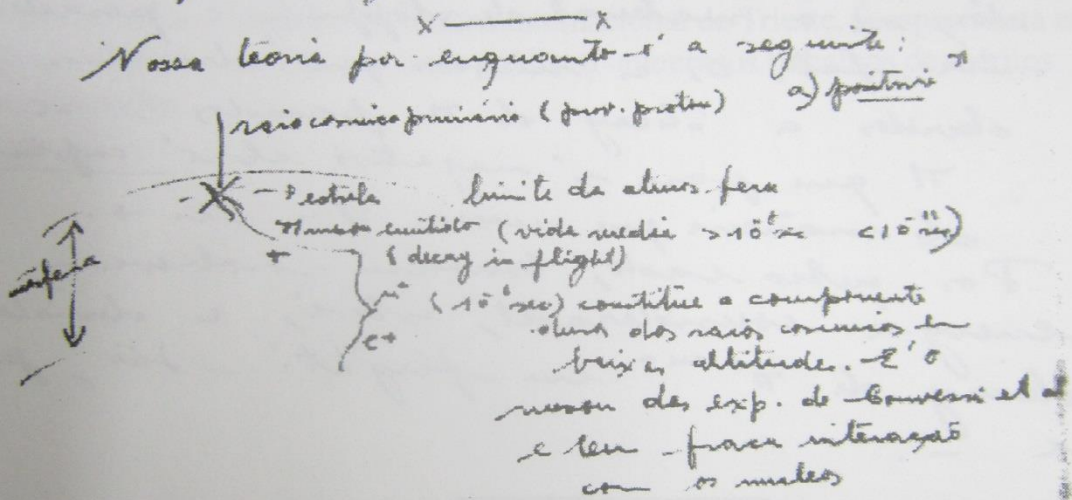


Consideração de ângulo sólido mostram  
 que a emissão de mesons por nucléos em explosão  
 (provocada por primários não identificados) é  
 muito comum. Além disso nós podemos  
 observar mesons de energia  $< 6.4\text{GeV}$ , de maneira  
 que, levando em conta o ângulo sólido e

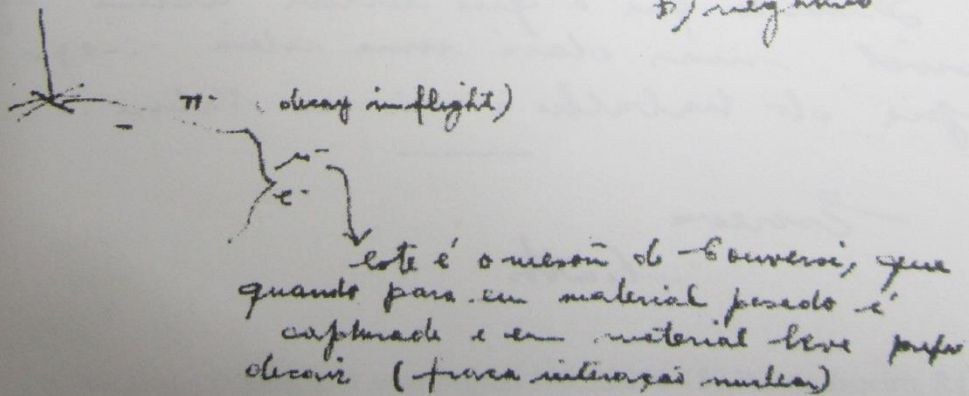
o depósito de energia (superfície, p. ex., superfície),  
 sempre, e que em 1ª aproximação, mesons  
 são emitidos por núcleos, sempre que haja  
 energia suficiente para isso.

Simple estudo mostra que os mesons  
 que observamos terminam na placa são produzi-  
 dos pelo material mesmo da placa. Não  
 há lugar para mesons<sup>cos</sup> próximos da atmosfera,  
 e que mostra que a vida é bem mais  
 curta que  $2 \times 10^{-6}$  seg ( $10^4$  cm).

Nossa teoria por enquanto é a seguinte:



b) negativo



c) negativo (2)



se o  $\pi$  mesmo para antes de  
desintegrar, ele é sempre  
capturado (parte interessante)

c) Explicar porque nossos  $\pi$  mesmo não são tão  
estúpidos apesar de serem  $+ + -$  (-convên),  
enquanto que -convên acha que  $\pi$  são captura-  
dos, e o material de stopping é pesado.  
De fato, os  $\pi$  mesmos que observamos são  
devidos a decay de  $\pi$  parados e se o  
 $\pi$  que para é negativo ele é capturado,  
de maneira que nossos  $\pi$  são  $+ +$ .  
Por outro lado, convên observam a ob-  
servação considerável, isto é, a devida a  
decay de  $\pi^{+ ou -}$  em flight. São, pois,  $+ +$   
e  $- -$ .

Dunio que o que escrevi acima seja compre-  
sível, mas clari como seria logo mandarei  
cópia do trabalho à sair em Nature.

Escreva

L.attes

[64] Personal History and application for a fellowship (Rockefeller Archive Center, Sleepy Hollow, 1946), número localizador 244, caixa 81, pasta 1541.

FUTURE PROSPECTS:

What are your plans for the future, and how definite are they? To stay in this laboratory,  
working on the application of the photographic technic to nuclear physics  
during at least one year. Before going back to Brasil, I intend to  
spend some months in another laboratory in Europe (Institute du Radium  
in Paris or Inst. of Theoret. Physics in Copenhagen). Have you a definite post to which to return at the expiration of a fellowship?

Yes



[76] G. Wataghin para E. Lawrence (The Bancroft Library, University of California, Berkeley, 1947), correspondência, rolo 15, caixa 10, pasta 35.

DEPARTAMENTO DE FÍSICA  
 FACULDADE DE FILOSOFIA, CIÊNCIAS E LETRAS  
 DA UNIVERSIDADE DE SÃO PAULO  
 AV. BRIGADEIRO LUIZ ANTONIO N.º 754  
 SÃO PAULO (BRASIL)

[ref. EOL:EI]

Copenhagen, Dec. 14<sup>th</sup> 1947

Dear Professor Lawrence,

I am writing to you from the institute of prof. Bohr, where Dr. Lattes and myself had the opportunity to discuss some of the recent experiments concerning the existence of several types of mesons (Lattes, Occhiattini & Powell) and the production of mesons by the primary radiation.

Before I left Brazil (on Dec. 8<sup>th</sup>) I received the official information from the Brazilian Embassy in Washington, that the Atomic Energy Commission has granted to Lattes the permission to work in your Laboratory. The Rockefeller Foundation received also this information. Lattes is returning for Christmas to Brazil (by airplane) and, after about of 1 month of rest, he will be ready to go to Berkeley, e.g. at the beginning of February 1948.

I want to thank you once more for the opportunity you give to Lattes to work in your laboratory and I shall be grateful if you could write to us when you would like him to start to work in Berkeley. I shall travel in Europe until

University of California - Berkeley  
 The Bancroft Library  
 Dr. Ennio Lawrence

DEPARTAMENTO DE FÍSICA  
 FACULDADE DE FILOSOFIA, CIÊNCIAS E LETRAS  
 DA UNIVERSIDADE DE SÃO PAULO

AV. BRIGADEIRO LUIZ ANTONIO N.º 754  
 SÃO PAULO (BRASIL)

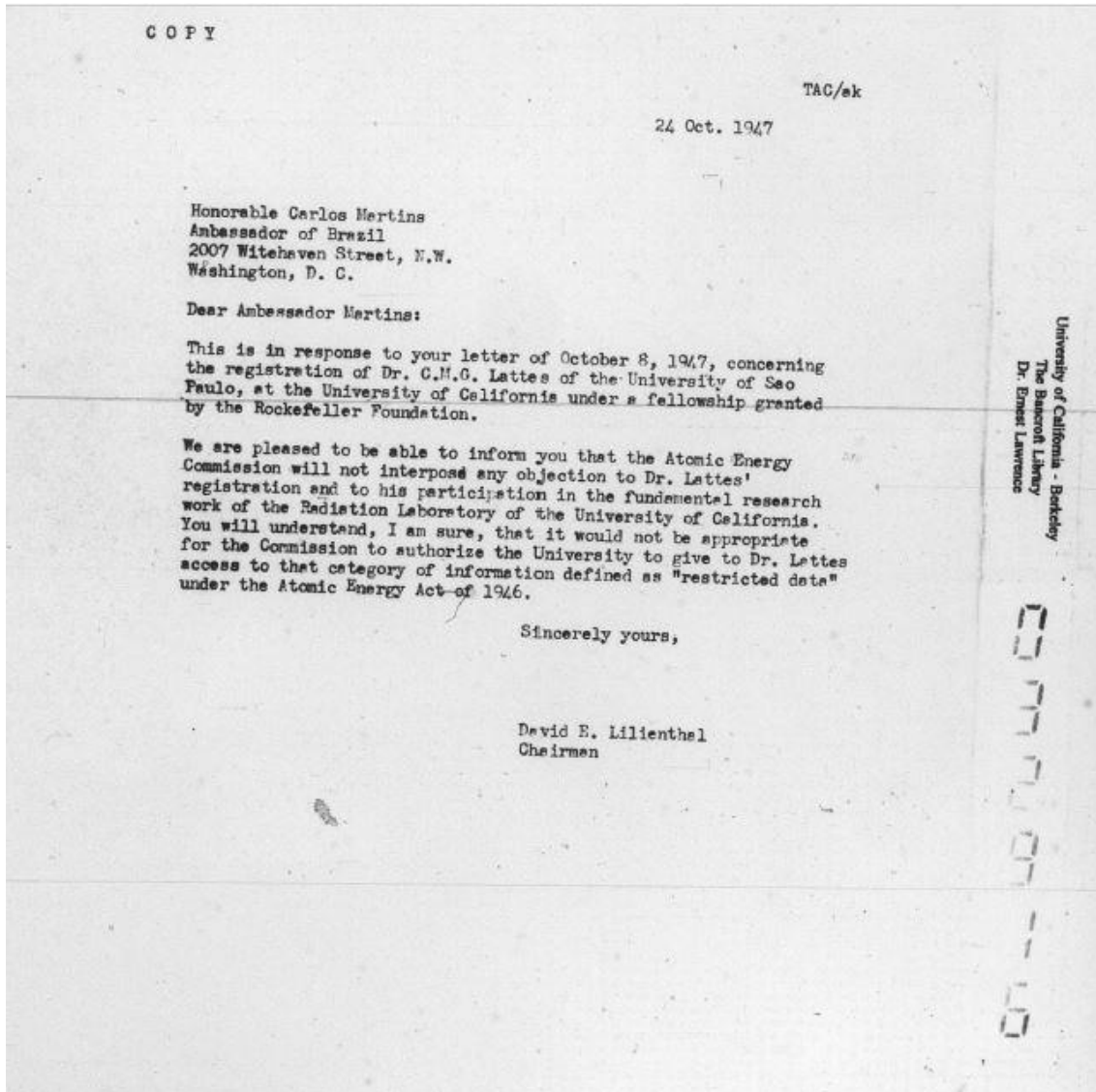
February 15<sup>th</sup>, so I take the liberty to  
 ask if you could write directly to C. Watters  
 at the address of our Department of physics in  
 S. Paulo. I took already the necessary steps  
 in order that the Brazilian authorities  
 (the Rector and the Dean of our Faculty) give  
 to Watters the permission to stay abroad for  
 more one year as a Fellow of the Rockefeller  
 Foundation.

Watters and myself are both really happy  
 about the possibilities of work he will have  
 in Berkeley, and we both are sending  
 to you our best wishes for  
 Merry Christmas and Happiest New Year.

Yours very cordially  
 Gilberto Wataglin

Professor Ernest O. Lawrence  
 Radiation Laboratory  
 Berkeley 4, California.

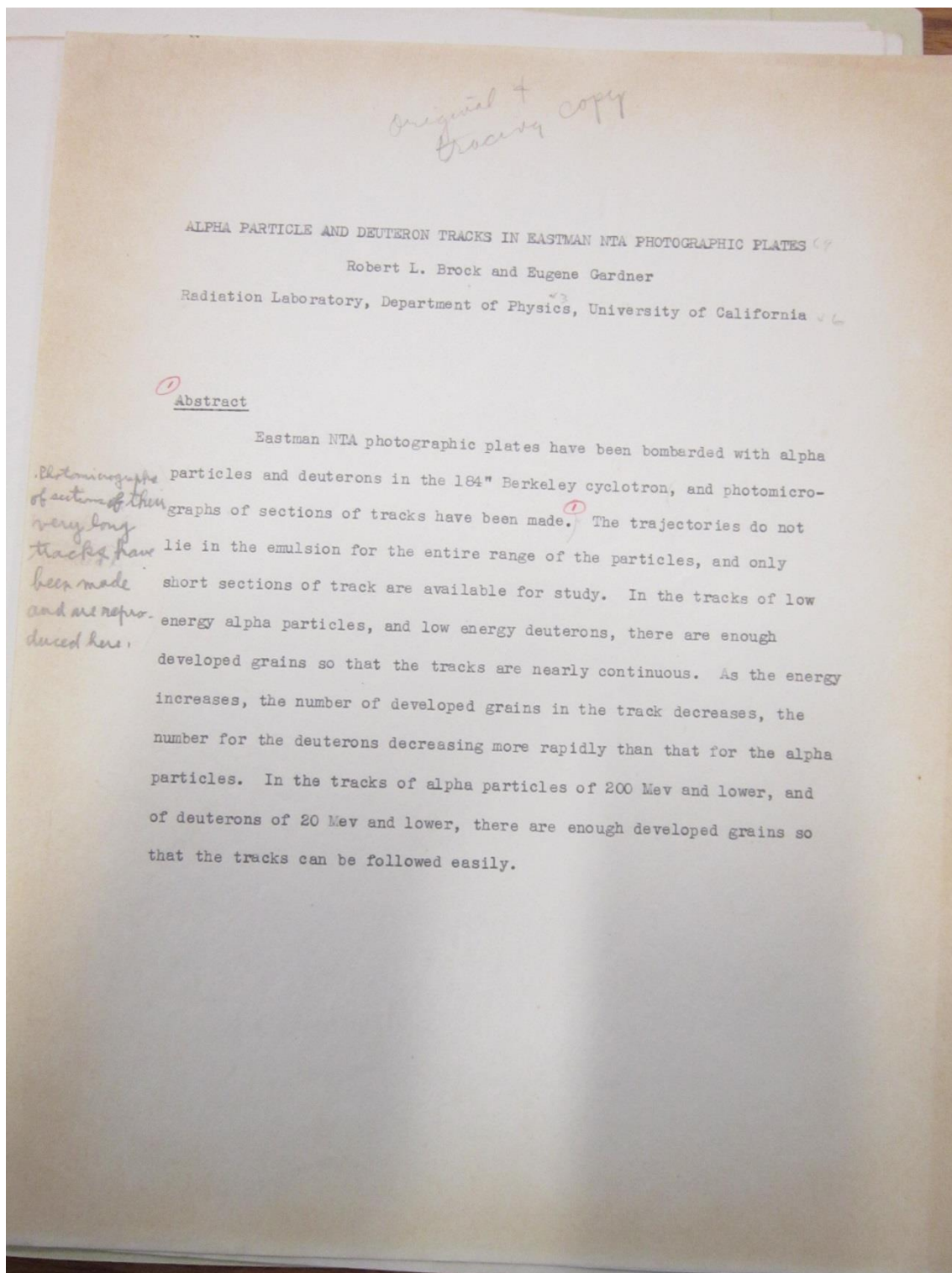
[77] D. Lilienthal para C. Martins (The Bancroft Library, University of California, Berkeley, 1947), correspondência, rolo 15, caixa 10, pasta 35.



[78] Status of the UCRL research work on July 09 (Logbooks of Meson detection experiments by Gardner Research Group, National Archives and Records Administration, San Bruno, 1947), caixa 1, pasta: 6 Photoemulsion method of meson detection.

PERSONNEL		DESCRIPTION OF PROJECT	STATUS OF WORK	STATUS OF PUBLICATION		
Reported By	Others			Stanford Meeting Papers	Letters to Editors	Journal Papers
STATUS OF UCRL RESEARCH WORK ON JULY 9, 1947, page 2						
Alvarez	Morley Fanozsky Richard, Cook Johnston	1. Half life of neutron 2. Proton - Proton scattering using film detection 3. Proton - Proton scattering using counters	Apparatus being built (Waiting completion of Linear Accelerator)			
McMillan	Helmholtz-Sewell Sewell, Fatierson Enax	1. Energy and half life of Be 10 2. Angular spread of 164 Mev neutron beam 3. Cross section measurements a) Total cross section for neutrons over 20 Mev in C, D, and H. (Will be continued for other elements) b) Relative cross sections for formation of C11, B11, O15 etc. in light elements with neutrons c) Absolute cross section for C12 (p, pn) C* reaction 4. Relative cross section of C12 (p, pn) C* and O12 (n, 2n) O11 to obtain absolute values for reactions of 3B above	Completed Completed Complete in 1 month. Starting Not started yet Not started	One		Sent to Phys. Rev. Declassified
Thornton	Stephan	Excitation vs. energy for C12 (D, 2n) C11 and O12 (s, an) O11 reactions with 60° and 164° Cyclotrons - approximate absolute values Range measurements - 200 Mev Deuterons	Almost complete Started			
Fowler		Cloud Chamber Observations in 164 Mev Neutron Beam 1. Proton Recoils a) Inside shield - total recoils from C and Paraffin b) Outside shield with collimated beam - recoils over 100 Mev 2. Stars - Qualitative observations	700 tracks obs. In Progress Good pictures obtained	One	Should be written	Paper Suggested
Moyer	Janney Farley Leith & Bratenahl Hildebrand, Enable, Farley & York Medical Group	1. Search for Negative Protons 1. C14 Instrumentation 2. Isotope Program a) Identification of Co activities b) Identification of Cu activities c) Calutron d) C14 Preparation for Jenkins 3. Health Physics Program a) Physical Dosage (ionization) in neutron beam b) Animal dosage in neutron beam	Continuing Complete Complete Complete Method developed In progress Barely started	One One	Published Ready	Should be written up for ESI or JAP Should be written Classified report, should be written
Seaborg		1. Radio-Isotopes of Po & Bi using 60° and 164° Cyclotrons 2. Fission of Bi 3. Radium Series 4. Transuramic Elements - including work done at Chicago 5. Uranium Fission with alpha particles 6. Radioactive yields from 200 Mev Deuterons and 400 Mev alphas in several areas below Bi (Sn As & O) not inc. fission. Address by Seaborg before fall meeting of American Chemical Society on high energy bombardments 7. Fission Cross Sections 8. Secondary reactions	Continuing In Progress 3 or 4 months work remaining 3 or 4 months work remaining Designing instruments Not active	Two	August Phys. Rev. August Phys. Rev.	2 to Phys. Rev. by one week To be written in 2 or 3 months for P.R. Being declassified for Chicago Reports Will be published
Segre	Wiegand Wiegand Zelly	EXPERIMENTAL PHYSICS 1. Pq Bi and Tl Fission 2. Stopping power of materials 3. Chemical properties of Astatine 4. Search for Be 7 5. Deuteron excitation of Bi - absolute cross sections	Should be completed in 4 weeks	One One One	Should be written immediately on Bi (unless held for Segre's return)	
STATUS OF UCRL RESEARCH WORK ON JULY 9, 1947, page 3						
PERSONNEL		DESCRIPTION OF PROJECT	STATUS OF WORK	STATUS OF PUBLICATION		
Reported By	Others			Stanford Meeting Papers	Letters to Editors	Journal Papers
Moyer (cont'd)	Hildebrand Enable Enable	4. Shielding effectiveness of materials a) Ionization chamber measurements b) Coincidence counting measurements c) Secondary neutron intensity and energy 5. Neutron scattering	Started Started	One	Could write note	
Serber	Bohm & Feldy Bohm & Feldy Henschel, Val. Sewell Serber Feldy Serber Staff Horning Serber	THEORETICAL PHYSICS 1. Synchrotron Theory 2. Synchro-Cyclotron Theory 3. 164 Mev Performance (with experimental results) 4. Linear Accelerator Theory 5. Proton Proton Scattering Calculations for E. Villan 6. Angular Distribution of Neutron Beam 7. Range curves 8. Star distribution 9. Shielding Synchrotron (electron absorption)	Complete Complete Complete Complete Complete Nearly complete One month Complete	Two		Published Sent to P.R. To be written Should be written up. Being written Will be written in one week Should be written up.

[79] *Alpha Particles and Deuterons Tracks in Eastman NTA Photographic Plates* (Logbooks of Meson detection experiments by Gardner Research Group, National Archives and Records Administration, San Bruno), caixa 1, pasta: Typescripts of publications and correspondence Robert L. Brock.



T.O.L.

ALPHA PARTICLE AND DEUTERON TRACKS IN EASTMAN NTA PHOTOGRAPHIC PLATES

Robert L. Brock\* and Eugene Gardner

Radiation Laboratory, Department of Physics, University of California

Photographic plates<sup>1</sup> have been bombarded with high energy alpha

<sup>1</sup>For references to work on the use of photographic plates as charged particle detectors up to 1941, see review article by Maurice M. Shapiro, R.M.P. 13, 58 (1941). For references to more recent work, see C. F. Powell and G. P. S. Occhialini, Nuclear Physics in Photographs (Clarendon Press, Oxford, 1947).

particles and deuterons in the 184" Berkeley cyclotron<sup>2</sup> for the purpose

<sup>2</sup>W. M. Brobeck, E. O. Lawrence, et al. Phys. Rev. 71, 449 (1947)

of studying stars in the emulsion produced by high energy particles. <sup>3, 4, 5</sup> Al-

though the primary objective of the program was the study of stars, the plates afforded an opportunity for observation of the alpha particle and deuteron tracks, and in this paper these tracks will be described in order to illustrate the applicability and to show the limitations of the photographic plate method of detecting high energy charged particles.

*Description of results*

*because the range is so much greater than the thickness*

*in various parts of their range*

<sup>4</sup>The plates are exposed in the cyclotron with the plane of the emulsion approximately tangent to the beam trajectories. When the developed plates are viewed under the microscope it is seen that some tracks enter from the edge of the emulsion, (and some come in from the top or bottom.

Most of the particles leave the emulsion before they stop, so that only sections of tracks are available. Although the sections of tracks can be followed through many fields of view under the microscope, the actual

*start start at and some start at the bottom showing that the particles have come from the glass support.*

\*Now at the University of Washington, Seattle, Washington

lengths are rather short, seldom over a few millimeters in length. If a section of track is close to the edge where the beam first strikes the emulsion, it is assumed that the energy of the particle causing the track is the same as the energy of the beam particles at that radius. By placing plates at various radii, it is possible to obtain bombardments by particles of any energy up to the maximum available from the cyclotron. This maximum energy is about 380 Mev for alpha particles and 190 Mev for deuterons.<sup>3</sup>

*Table I shows the ranges for various energies of alpha rays and deuterons (which may amount to as much as several mm of emulsion)*

<sup>3</sup> Walter Stephan and R. L. Thornton, To be published

It is possible to find tracks which stop in the emulsion, but when these tracks are followed in the direction of higher energy, it is found that most of them leave the emulsion within a millimeter or so of the low energy end. Sometimes the high energy ends of deuteron tracks are lost in the general background of developed grains before they leave the emulsion. A group of six alpha particle tracks and six deuteron tracks whose low energy ends lie in the emulsion are shown in Fig. 1. The tracks have been arranged so that their low energy ends are at the zero on the scale. Tracks of length greater than 0.8 mm are shown in two sections, both the tracks and the scale being broken at the 0.8 mm point. The plate from which these tracks were taken was exposed at a position in the cyclotron corresponding to 20 per cent of the full energy available. The particle trajectories did not, however, lie in the emulsion for their entire range, and the sections of tracks shown in Fig. 1 correspond to energies of the order of 60 Mev for the alpha particles and 20 Mev for the deuterons. Sections of some of the tracks shown in Fig. 1 are shown on a larger scale in Fig. 2. The sections shown in Fig. 2 are at a distance of about 0.4 mm from the low energy ends of the tracks.

*Table I*  

Energy	Range
alpha	mm
d	mm
d	mm

*Range*

*Grain densities* (5)

The grain densities (i.e., number of grains per unit length) of the tracks shown in Fig. 1 have been plotted in Fig. 3. On the basis of studies

-3-

of this kind, it is often possible to distinguish between the two types of tracks, since the points corresponding to the alpha particle tracks lie above those for the deuteron tracks. For long tracks which stop in the emulsion, it is easy to distinguish between alpha particle and deuteron tracks. If a track does not stop in the emulsion, however, one does not know where on the abscissa scale the grain density values are to be plotted, and it may not be possible to identify the track, particularly if only a small section of track is available. <sup>(6)</sup> In studies which involve tracks in random directions (e.g., studies of particles from stars in photographic emulsions) a large fraction of the sections of track observed may be too short to enable one to identify the particle causing the track. <sup>(6)</sup>

It might be expected that the grain density as shown in Fig. 3 would be proportional to the rate of energy loss of the particle causing the track, and this is probably true for portions of tracks where the grain density is ~~very~~ low. As the rate of energy loss becomes larger and larger, more and more grains in the path are rendered developable, but this cannot go on indefinitely because there are only a fixed number of grains along the path, and when all of these grains have been rendered developable, the grain density cannot increase any more no matter how large the rate of energy loss becomes. Thus, for portions of the track where the grains are close together, the grain density is not a good measure of the rate of energy loss.

Maximum  
energy recorded. <sup>(7)</sup>

In the low energy ends of both alpha particle and deuteron tracks there are enough developed grains so that the tracks are nearly continuous, as can be seen in Fig. 1. As the energy increases, the number of developed grains in the track decreases, as shown in Fig. 3. At very high energies, there are so few developed grains that the tracks become lost in the general background of developed grains. It is hard to give a definite value for the



-4-

energy above which tracks cannot be followed, because grain density required for following a track depends to such a great extent on the background of developed grains. If a plate is given a heavy exposure (i.e., is exposed to a large number of high energy particles) there may be so many developed grains on the plate that no individual tracks can be seen. On the other hand, if the plate has been given a very light exposure, it may be possible to follow tracks which consist of a very few developed grains. We have, at times, seen tracks on plates exposed to alpha particles at the full energy of the cyclotron, but this is only possible when the exposure is extremely light. <sup>(7)</sup> For tracks of very small grain density, local <sup>fluctuations</sup> in the grain density become important, and the eye will catch those tracks which have a few more grains than the average, and miss those which have a few less. Thus it is often possible to see tracks when it is not possible to count them accurately. In the tracks of alpha particles of 200 Mev and lower, and of deuterons of 20 Mev and lower, the tracks are continuous enough so that they can be followed easily, and they can be counted without much danger of missing any. Sections of tracks found in a plate exposed to 200 Mev alpha particles are shown in Fig. 4. In this figure, several alpha particle tracks are seen in the same field of view.

<sup>(8)</sup> All of the tracks shown in Fig. 1 were taken from the same plate, and the tracks shown in Fig. 4 were taken from a plate of the same emulsion number (i.e., same manufacturer's batch). Although we have exposed several hundred Eastman NTA plates on the cyclotron, nearly all of them have been of this same emulsion number. We are informed by Eastman Kodak Company that for NTA plates, as well as for other photographic materials, it is recommended that experimenters using these plates for quantitative work calibrate each emulsion of different emulsion number that they use. <sup>(9)</sup>

as regards stopping time: maximum energy recorded?

Ackno Identified plates (9)

Ilford Nuclear Research plates<sup>7</sup> are being used in experiments on

<sup>2</sup> Powell, Occhialini, Livesey and Chilton, J. Sci. Instr. 23, 102 (1946)

the 184" cyclotron, and we have had opportunity to observe alpha particle and deuteron tracks in various types of these plates. In Ilford B.1 and C.2 plates we have been able to follow deuteron tracks at higher energies than in Eastman NTA plates, but not at energies approaching the maximum available from the cyclotron. The work with these plates, as well as the work with the Eastman NTA plates was done in the spring of 1947, and since that time new types of plates have been developed. The Ilford type B.2 has superseded type B.1, and the Eastman NTB plates have been added to the NTA plates. We have not had enough experience with these new types of plates to be able to comment on them at this time. However, the manufacturers are aware of the need for plates which will record high energy deuterons, and it is to be expected that more sensitive plates have been, or will be, developed. <sup>(10)</sup> These more sensitive plates will not be suitable for distinguishing between alpha particle and deuteron tracks, because both low energy alpha particles and low energy deuterons will render developable practically all of the grains in the track. <sup>(10)</sup>

probably  
from grains

Proton tracks. (11)

No experimental data on proton tracks has been presented in this paper. It was easier to study alpha particles and deuterons, since the 184" cyclotron is arranged to accelerate these particles. Information obtained from the deuterons can be applied to proton tracks, since protons and deuterons of the same velocity produce very closely the same grain density. <sup>8</sup>

Still a comparison of the mean curvature of the tracks, which is greater for the deuteron, <sup>8</sup>

T. R. Wilkins and H. J. St. Helens, Phys. Rev. 54, 783 (1938)

may allow the distinction to be made. See fig 1

-7-

-6-

Thus we would expect the tracks of 10 Mev protons to have about the same grain density as the tracks of 20 Mev deuterons.

Acknowledgments (12)

The authors wish to express their appreciation to Professor Ernest O. Lawrence for his continued interest in this problem. We are indebted to Professor R. L. Thornton for helpful discussions. We also wish to thank Mr. A. J. Oliver for photographic work, Mr. Duane Sewell and the cyclotron crew for bombarding the plates, and the Eastman Kodak Company for information regarding Eastman NTA plates.

This paper is based on work performed under contract W-7405-Eng-48, with the Atomic Energy Commission, in connection with the Radiation Laboratory, University of California, Berkeley, California.

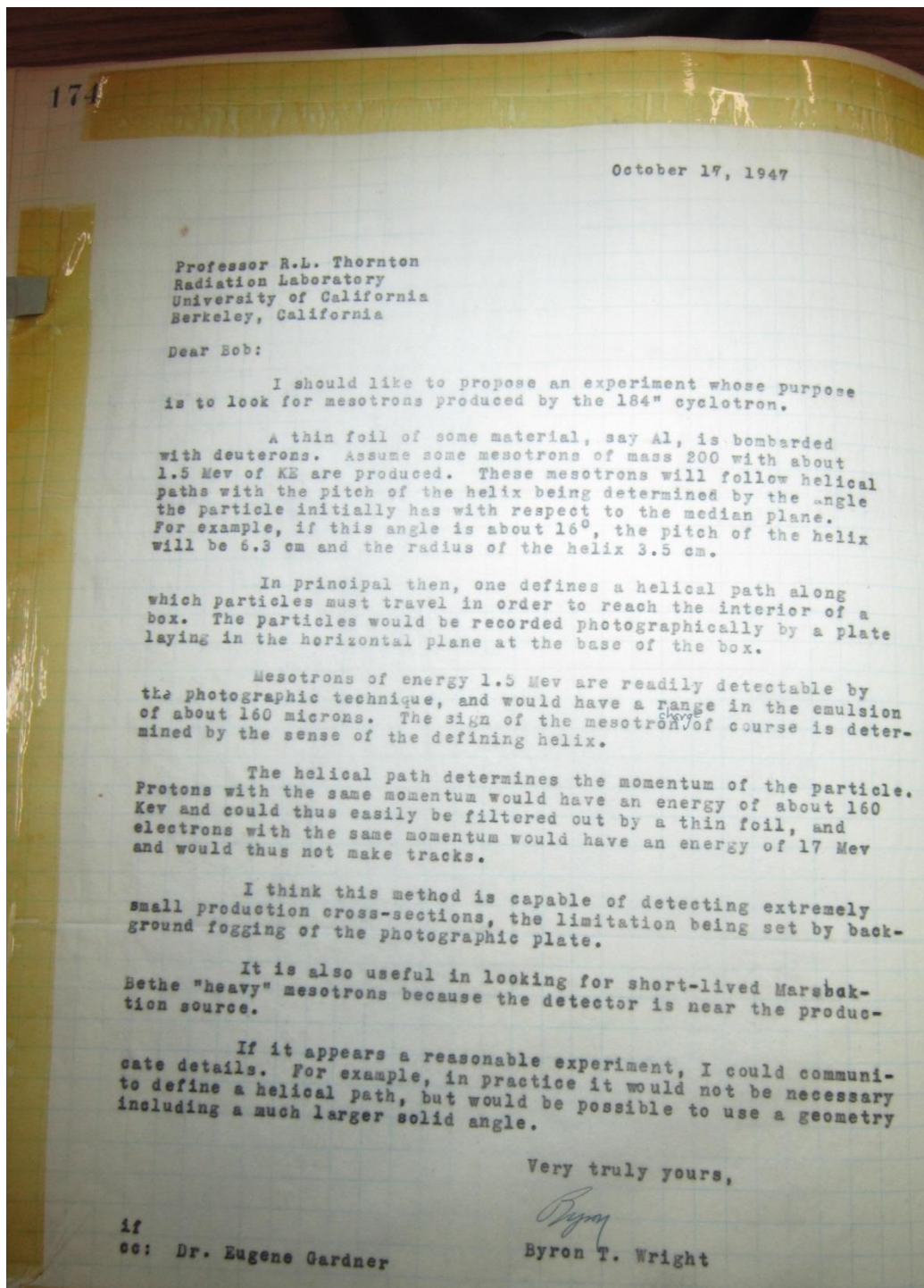
-7-

- Fig. 1.- Alpha particle and deuteron tracks formed in Eastman NTA photographic plate. (3-mm apochromatic objective lens.)
- Fig. 2.- Enlarged sections of tracks shown in Fig. 1.
- Fig. 3.- Grain densities of tracks shown in Fig. 1.
- Fig. 4.- Sections of tracks taken from Eastman NTA plate exposed to 200 Mev alpha particles. (3 mm apochromatic objective lens.)

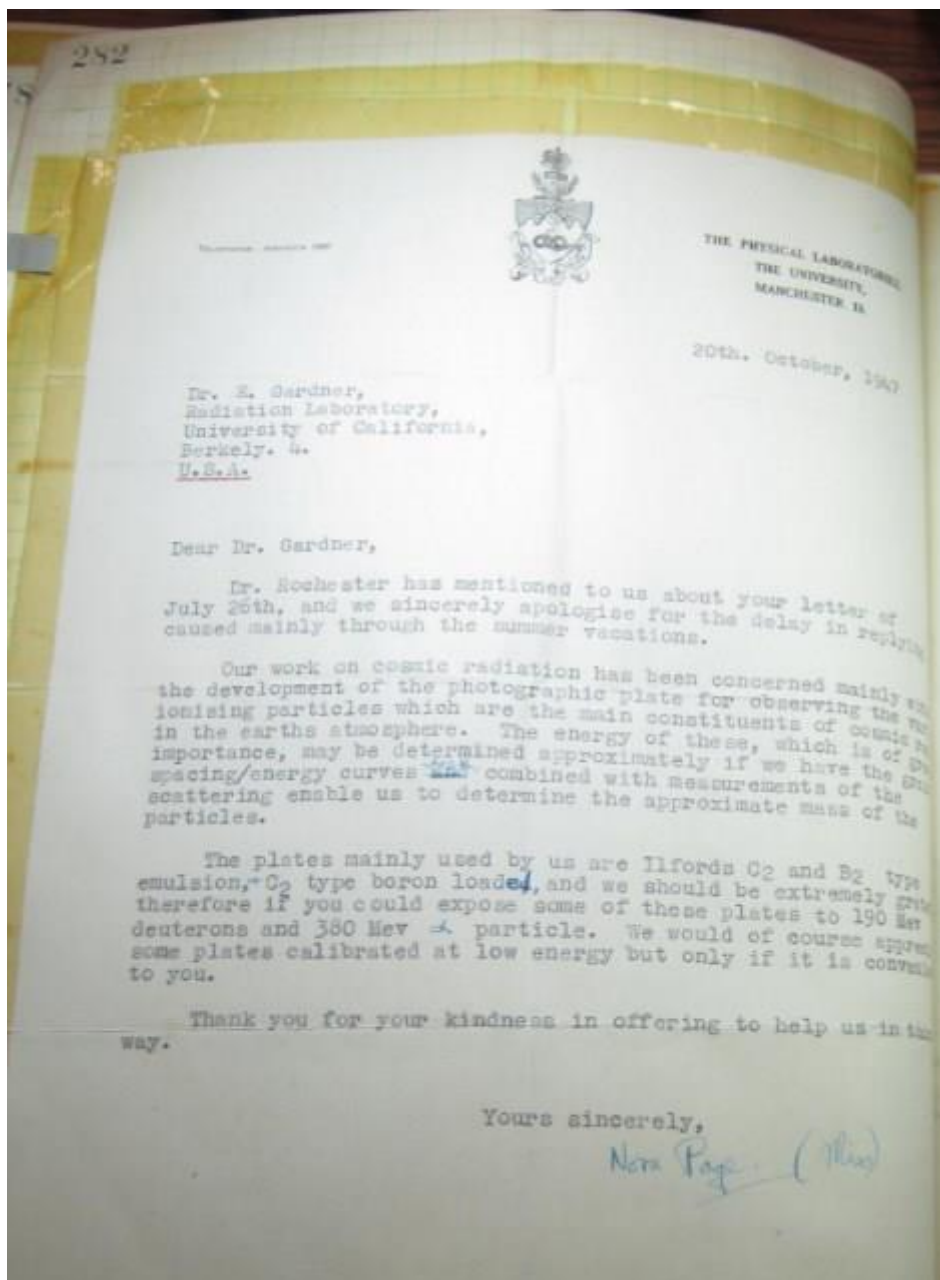
2.65 2.50 after  
 assembled from 250 prints from 150 negatives  
 of photomicrographs. The grain density decreases from end (0.0 mm)  
 to beginning. The grain densities are greater, and the tracks are more nearly straight  
 for a ray.

Would the authors care to draw a representative curve through  
 each series of points?

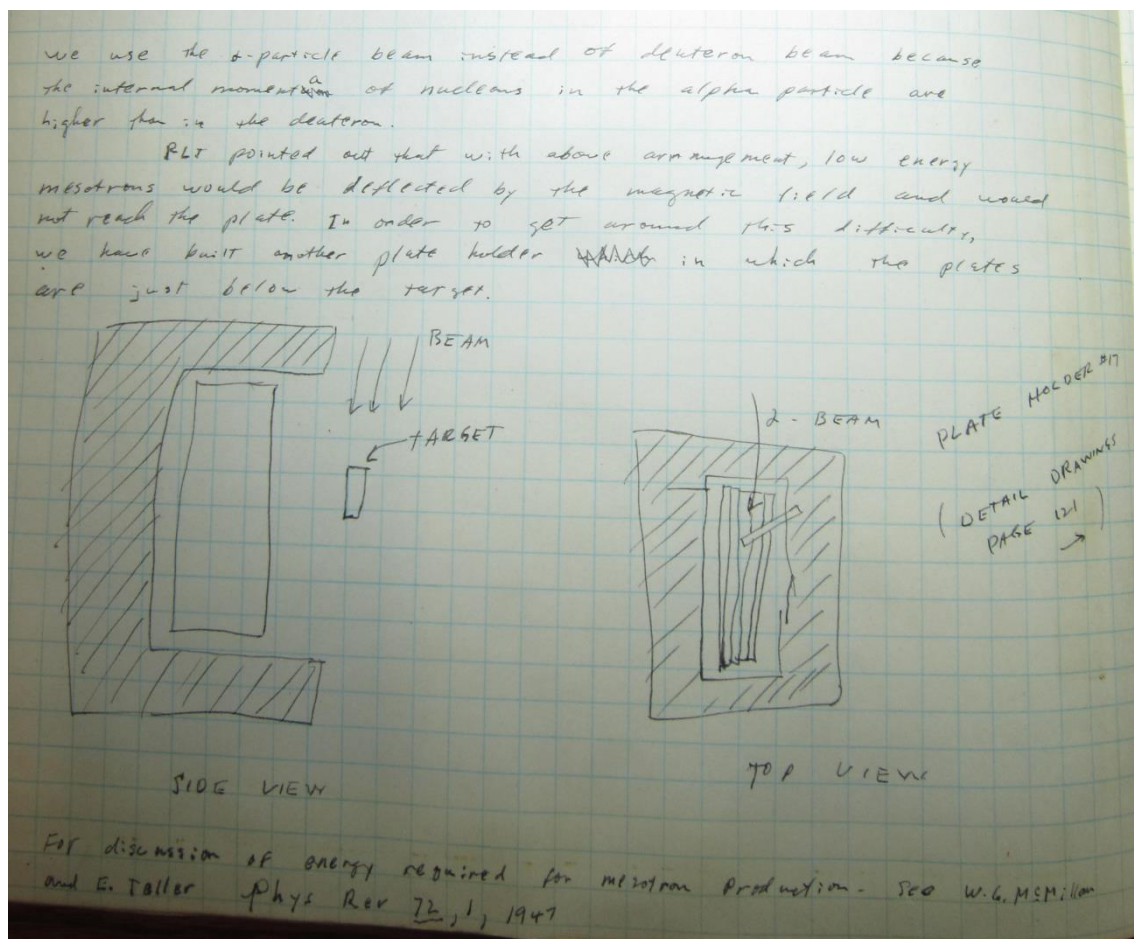
[80] B. Wright para R. Thornton (Logbooks of Meson detection experiments by Gardner Research Group, National Archives and Records Administration, San Bruno, 1947), correspondência, caixa 5, livro 09, página 174.



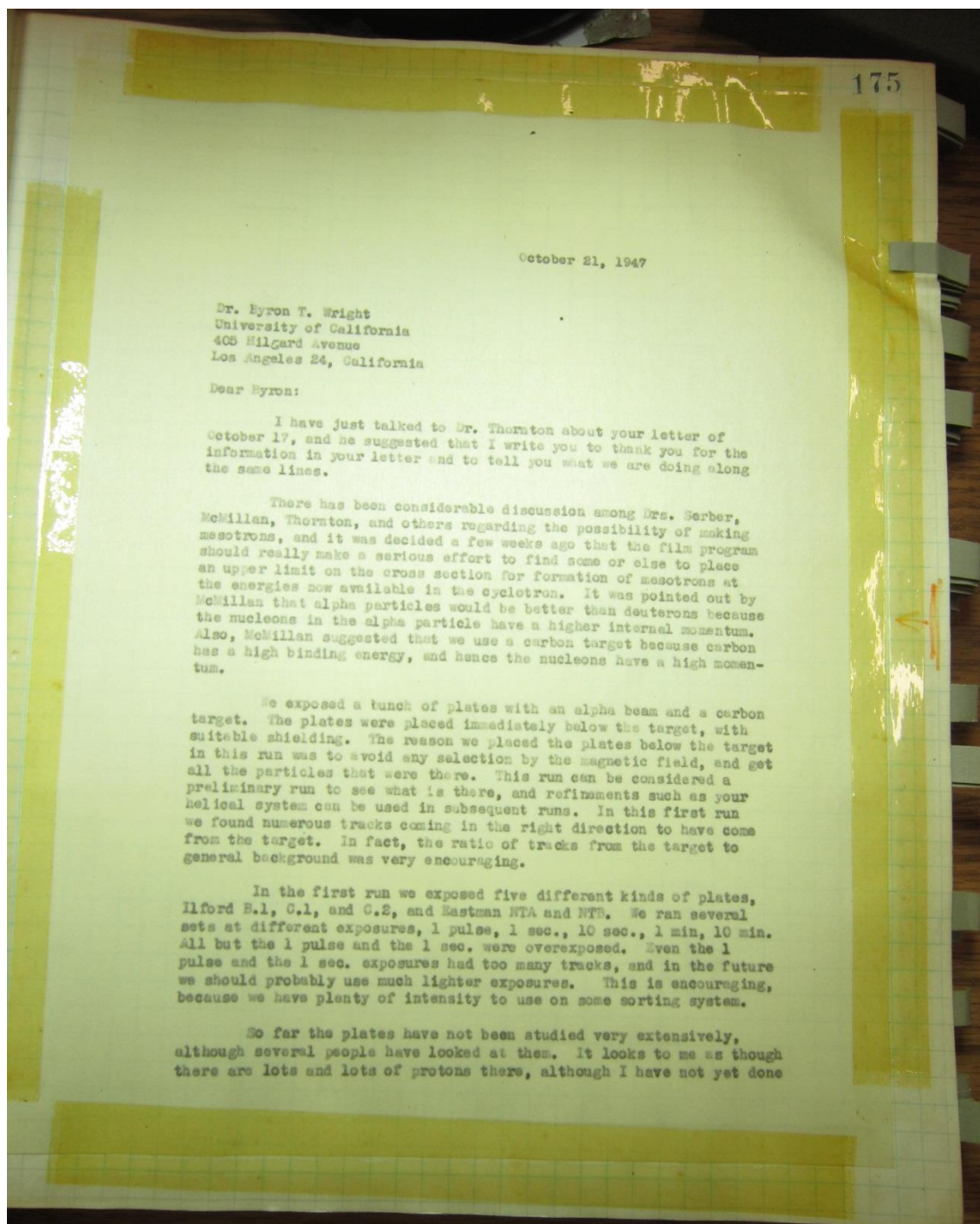
[81] N. Page para E. Gardner (Logbooks of Meson detection experiments by Gardner Research Group, National Archives and Records Administration, San Bruno, 1947), correspondência, caixa 5, livro 09, página 282.



[82] [Arranjo do aparato sugerido por E. McMillan, desenhado por E. Gardner, para tentar capturar mésons] (Logbooks of Meson detection experiments by Gardner Research Group, National Archives and Records Administration, San Bruno, 1947), caixa 5, livro 9, página 120.



[83] E. Gardner para B. Wright (Logbooks of Meson detection experiments by Gardner Research Group, National Archives and Records Administration, San Bruno, 1947), correspondência, caixa 5, livro 09, página 175.





Page Two  
Dr. Byron T. Wright  
Los Angeles, California

October 21, 1947

any grain count to tell for sure. The exposures on these plates are really too heavy to make them very good for study, and I think that we will run some more plates before doing any very serious looking.

Bob Thornton has worked out some of the details of helix trajectories in connection with proton-proton scattering. I think that they decided not to use them in proton-proton scattering, but his calculations might be a good start on calculation of some helix trajectories for mesotron study.

We have just sent a paper on grain density to the Review of Scientific Instruments for publication. I am sending you a copy for your files.

Very best regards to Reg and Ken, and to Lorna.

Yours truly,

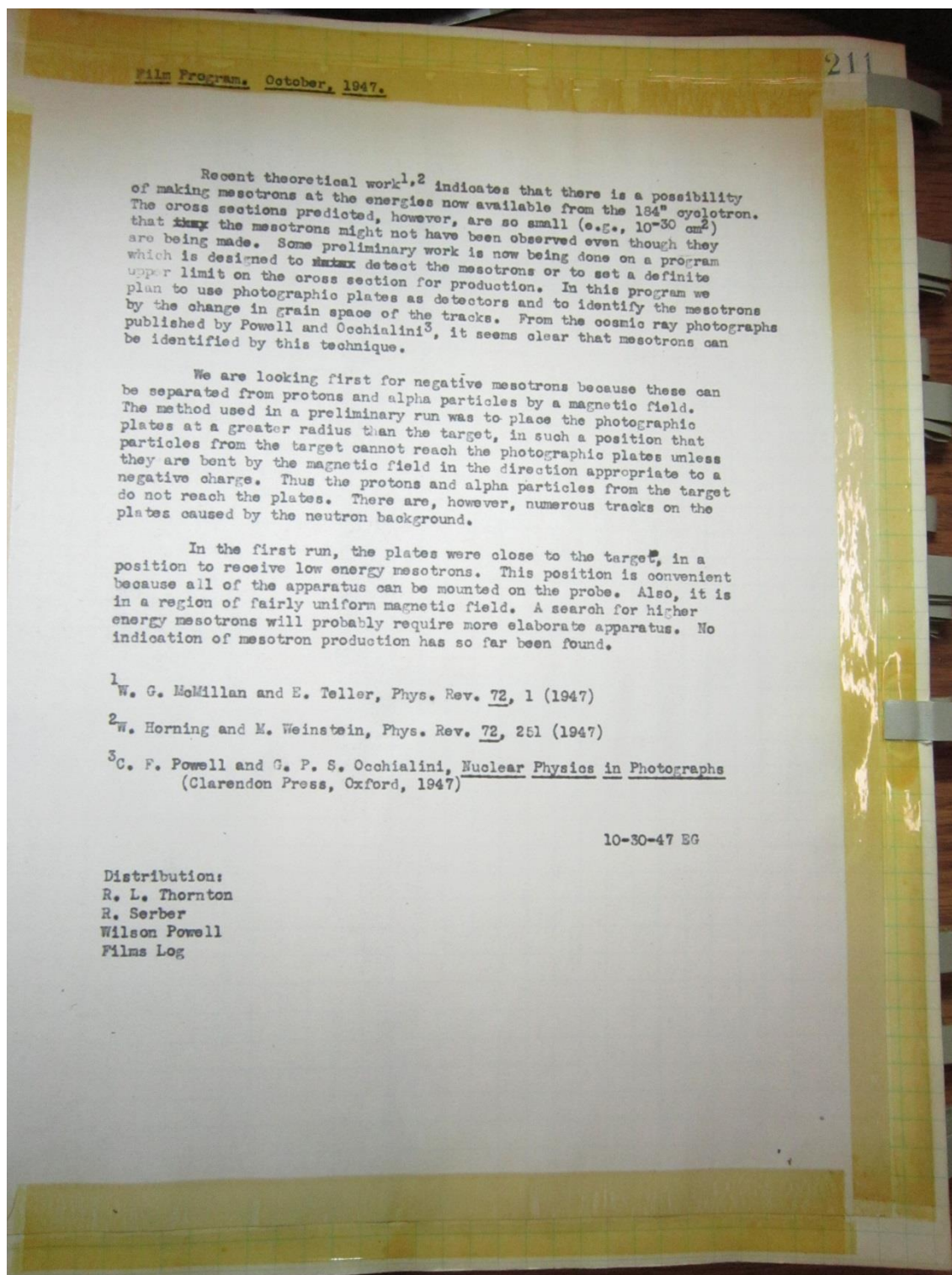
EG:ek

Eugene Gardner

Enclosure

cc: R. L. Thornton  
E. M. McMillan  
R. Serber

[84] Film Program (Logbooks of Meson detection experiments by Gardner Research Group, National Archives and Records Administration, San Bruno, 1947), caixa 5, livro 09, página 211.



Film Program. October, 1947.

211

Recent theoretical work<sup>1,2</sup> indicates that there is a possibility of making mesotrons at the energies now available from the 184" cyclotron. The cross sections predicted, however, are so small (e.g.,  $10^{-30}$  cm<sup>2</sup>) that ~~it is~~ the mesotrons might not have been observed even though they are being made. Some preliminary work is now being done on a program which is designed to ~~detect~~ detect the mesotrons or to set a definite upper limit on the cross section for production. In this program we plan to use photographic plates as detectors and to identify the mesotrons by the change in grain space of the tracks. From the cosmic ray photographs published by Powell and Occhialini<sup>3</sup>, it seems clear that mesotrons can be identified by this technique.

We are looking first for negative mesotrons because these can be separated from protons and alpha particles by a magnetic field. The method used in a preliminary run was to place the photographic plates at a greater radius than the target, in such a position that particles from the target cannot reach the photographic plates unless they are bent by the magnetic field in the direction appropriate to a negative charge. Thus the protons and alpha particles from the target do not reach the plates. There are, however, numerous tracks on the plates caused by the neutron background.

In the first run, the plates were close to the target<sup>4</sup>, in a position to receive low energy mesotrons. This position is convenient because all of the apparatus can be mounted on the probe. Also, it is in a region of fairly uniform magnetic field. A search for higher energy mesotrons will probably require more elaborate apparatus. No indication of mesotron production has so far been found.

<sup>1</sup>W. G. McMillan and E. Teller, Phys. Rev. 72, 1 (1947)

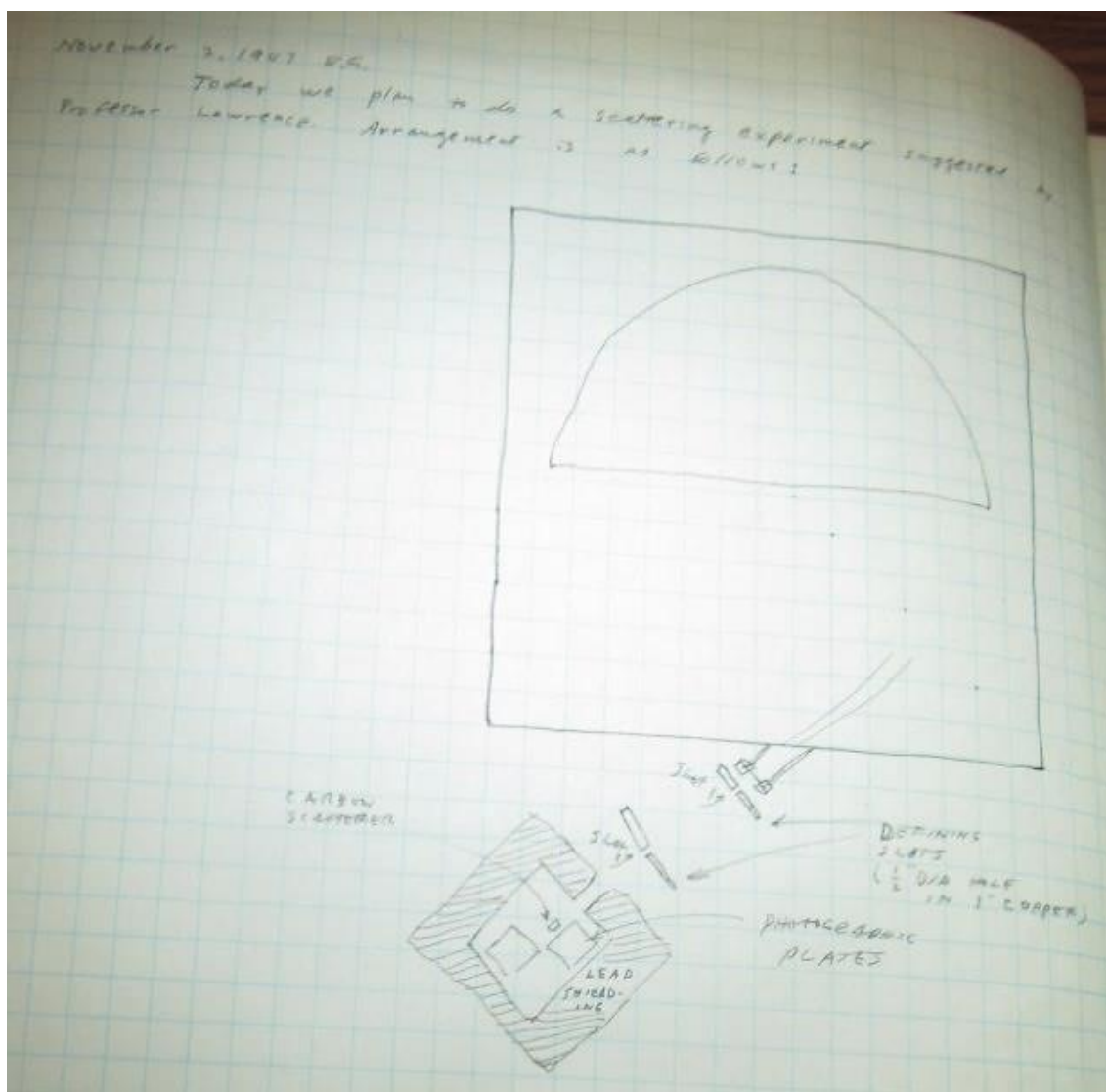
<sup>2</sup>W. Horning and M. Weinstein, Phys. Rev. 72, 251 (1947)

<sup>3</sup>C. F. Powell and G. P. S. Occhialini, Nuclear Physics in Photographs (Clarendon Press, Oxford, 1947)

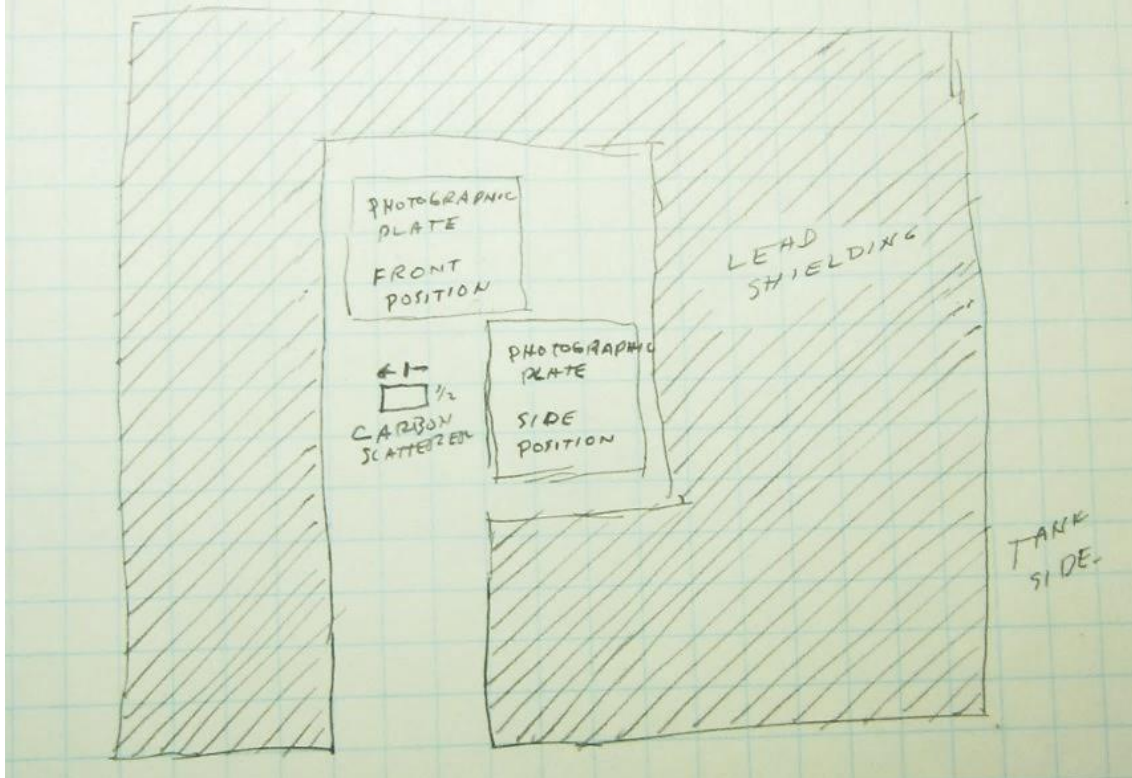
10-30-47 EG

Distribution:  
R. L. Thornton  
R. Serber  
Wilson Powell  
Films Log

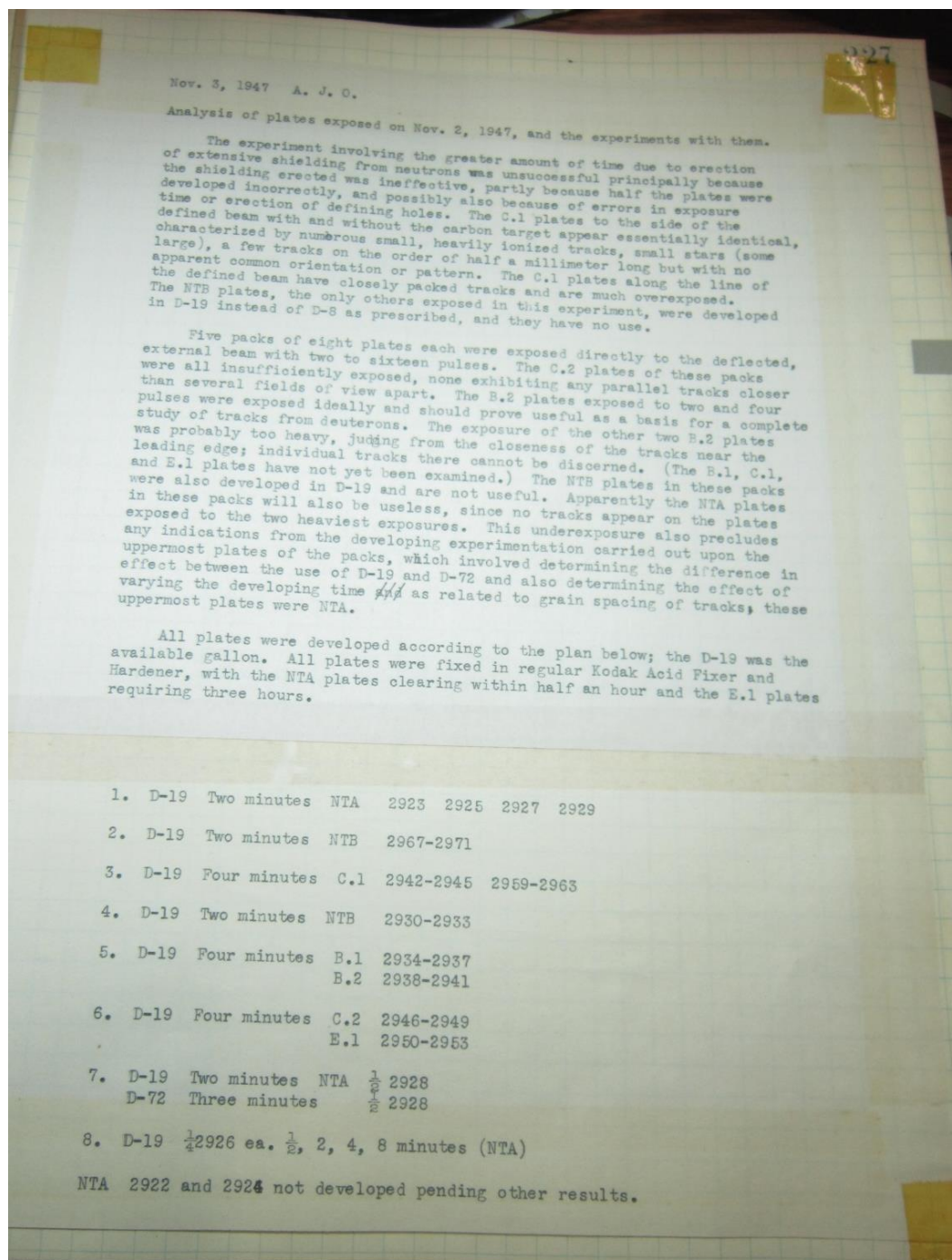
[85] [Experimento sugerido por Ernest Lawrence] (Logbooks of Meson detection experiments by Gardner Research Group, National Archives and Records Administration, San Bruno, 1947), caixa 5, livro 09, página 221.



Detail of lead shielding follows:



[86] Analysis of plates exposed on Nov. 03, 1947, and the experiment with them (Logbooks of Meson detection experiments by Gardner Research Group, National Archives and Records Administration, San Bruno, 1947), caixa 5, livro 09, página 227.



[87] [Nota escrita à mão por Eugene Gardner sobre os procedimentos adotados por Cecil Powell para revelar emulsões fotográficas] (Logbooks of Meson detection experiments by Gardner Research Group, National Archives and Records Administration, San Bruno, 1947), caixa 6, livro 10, página 95.

C. F. POWELL'S DEVELOPER  
(SEE NUCLEONIC PARTICLES IN PHOTOGRAPHY  
APPENDIX 4)

D-19						
H <sub>2</sub> O	500 cc					
Elon	2.2 g.	2.2	(Metol)	5.5 g		D-19 has equal amt of Elon (same thing)
Na Sulfite, des.	96.0 g	72.0		180 g		Powell's 75% of that in D-19
Hydroquinone	8.8 g.	5.8		22		same
Na CO <sub>3</sub> des. (carbonate)	48.0 g.	48.0		120		same
K Br	5.0 g.	4.0		10		80% of that in D-19
H <sub>2</sub> O to make	1000 cc	1000		2500		
no dilution						
4 min (max. var)						

Iod. sulfite: preservative when used  
in developer

Iod. carbonate: the surplus alkali  
or accelerator is dissolved  
removed

KBr: limits development to unexposed  
grains

Powell's developer differs only in that it has 25% less sod. sulfite (preservative for the developer) and 20% less pot. bromide than D-19 and he uses it diluted 1:4 water, developer for 33 min instead of 4

[88] Film Program (Logbooks of Meson detection experiments by Gardner Research Group, National Archives and Records Administration, San Bruno, 1947), caixa 6, livro 10, página 88.

Film Program. December, 1947

Search for Mesotrons

Work has continued on the search for mesotrons produced in the 184" cyclotron using photographic plates as detectors (see Progress Report, Film Program, October, 1947). The installation of the proton tube in the cyclotron makes available protons of energy considerably higher than the 95 Mev per nucleon previously available in the alpha particle beam. One experiment was done with a 1/16" carbon target placed in position to receive incident protons of energy 160 Mev. Photographic plates were placed at various radii on both sides of the target in order to record either positive or negative mesotrons over a considerable energy range. The plates show a rather high neutron background. The equipment for use in the proton tube is being redesigned in an effort to reduce the neutron background. No evidence of mesotron production has yet been found.

12-23-47 JWB

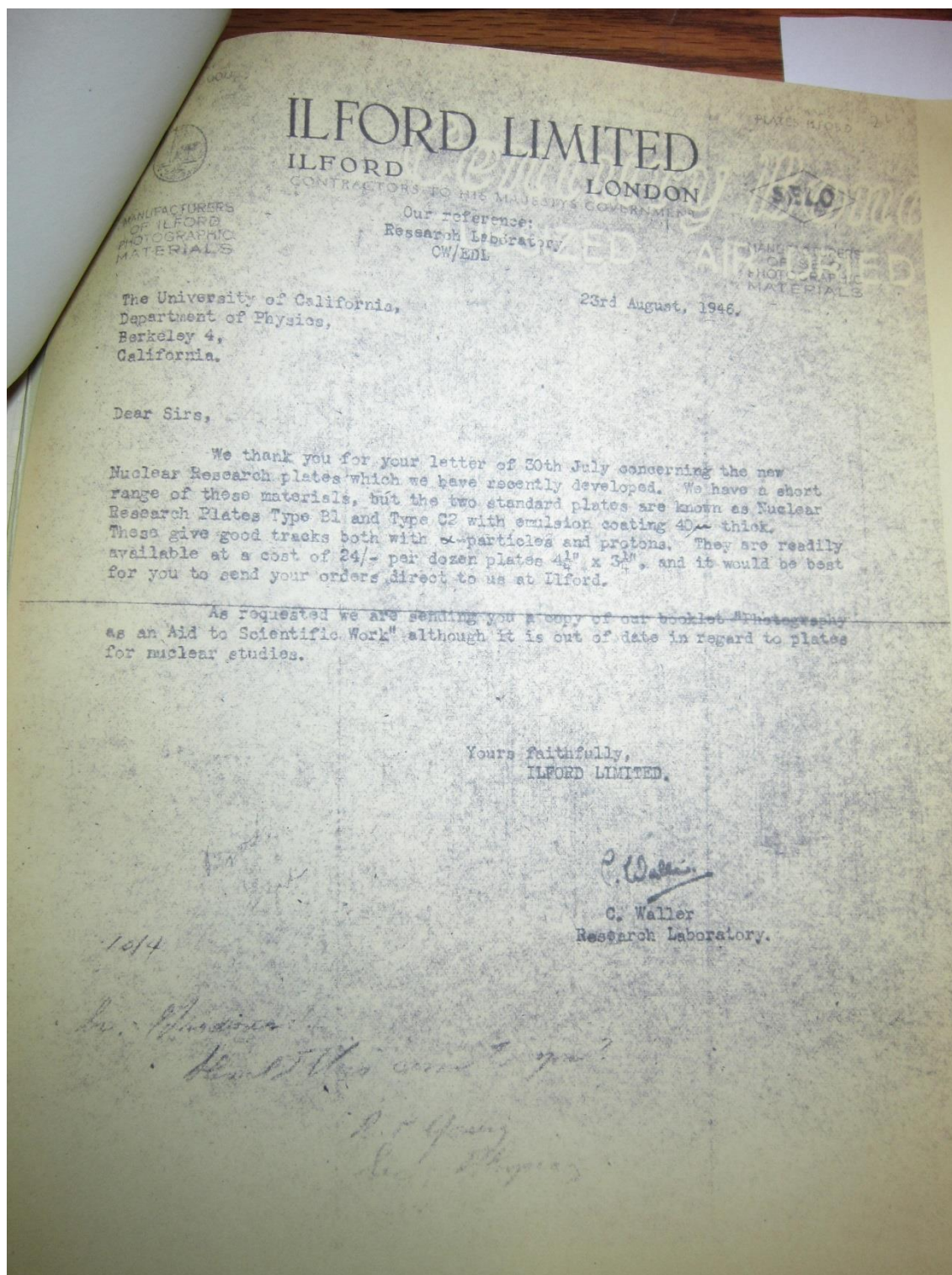
Scattering of Deuterons

Exploratory runs have continued in a study of the scattering of deuterons on lead and other materials. The first work was done with the circulating beam inside the tank, but more recent work with the deflected deuteron beam has been so much more promising that scattering in the circulating beam has been abandoned. The deflected deuteron beam has been brought out through the concrete shielding, so that the scattering experiment can now be done at a position where the neutron background is low. A scatterer is placed in the deuteron beam, and photographic plates are placed in positions to receive scattered particles in various directions. From the directions of the tracks observed on the plates, it is possible to tell which particles came from the vicinity of the scatterer. So far the experiments have been done in air, but we now plan to put the apparatus in a vacuum chamber.

12-31-47 EG

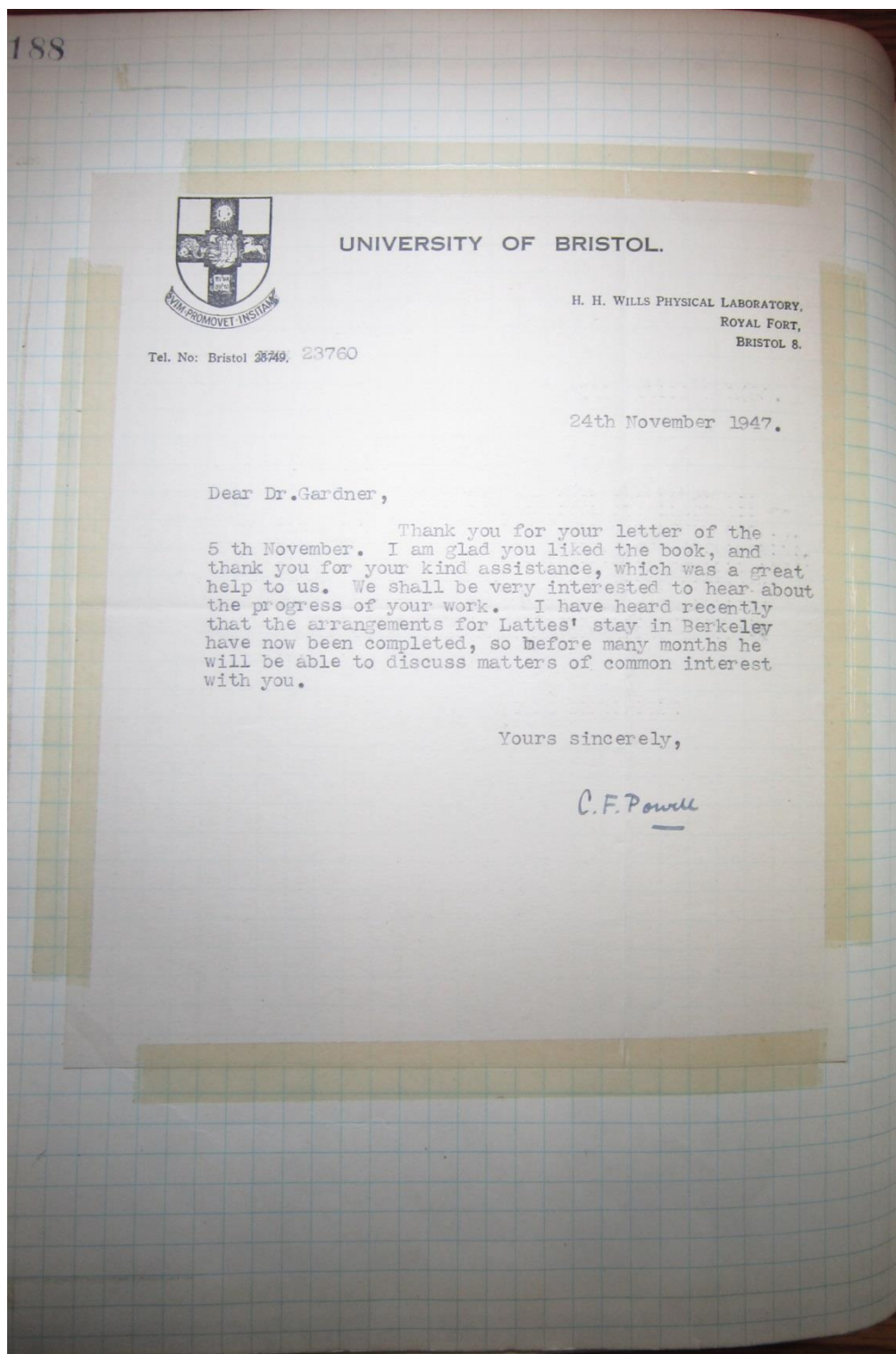
Distribution:  
R. L. Thornton  
R. Serber  
Wilson Powell  
Films Log

[89] C. Waller para Departamento de Física da Universidade de Berkeley (Research and Development Records and Administrative Files of Eugene Gardner, National Archives and Records Administration, San Bruno, 1946), correspondência, caixa 2, pasta 26.

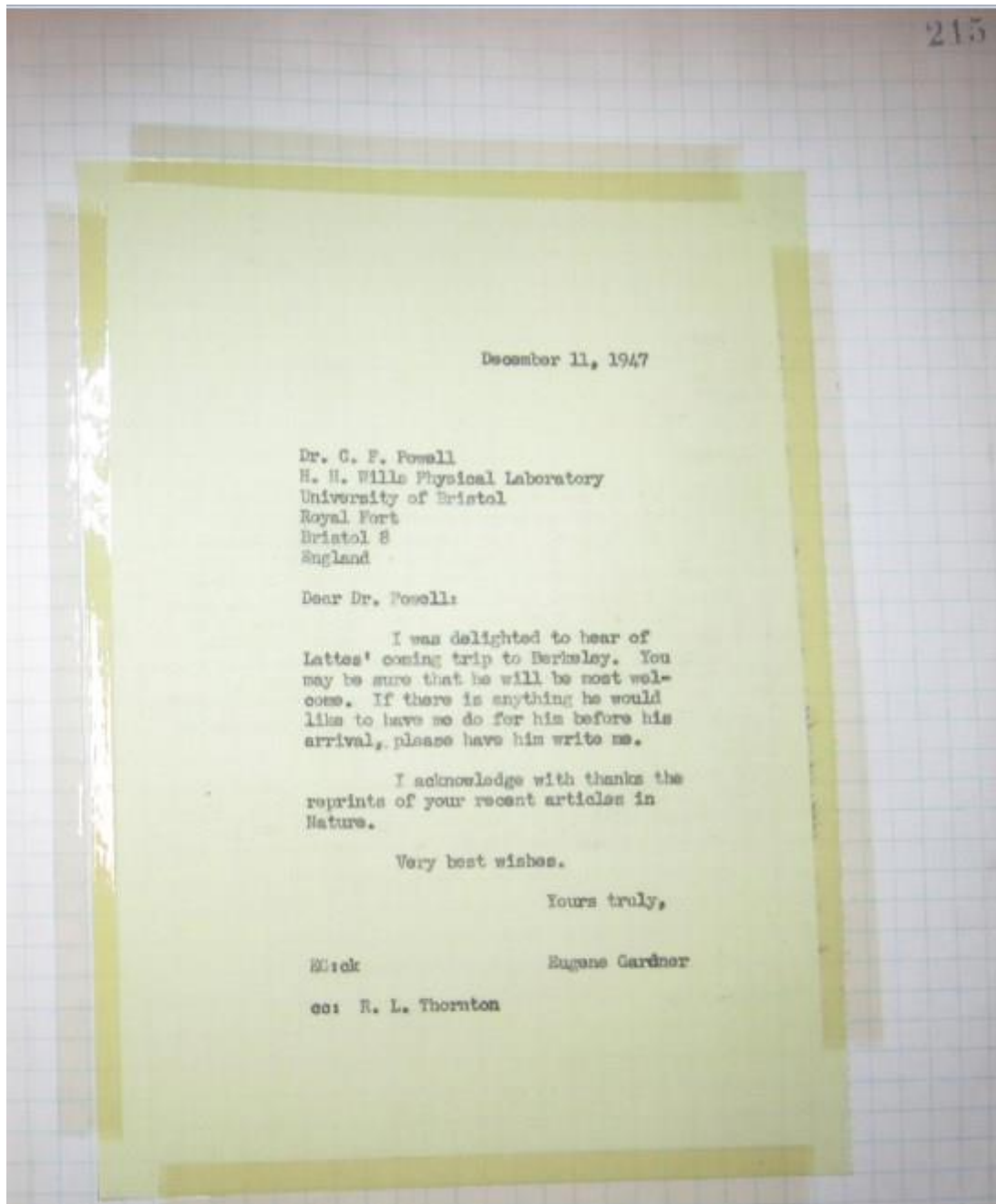




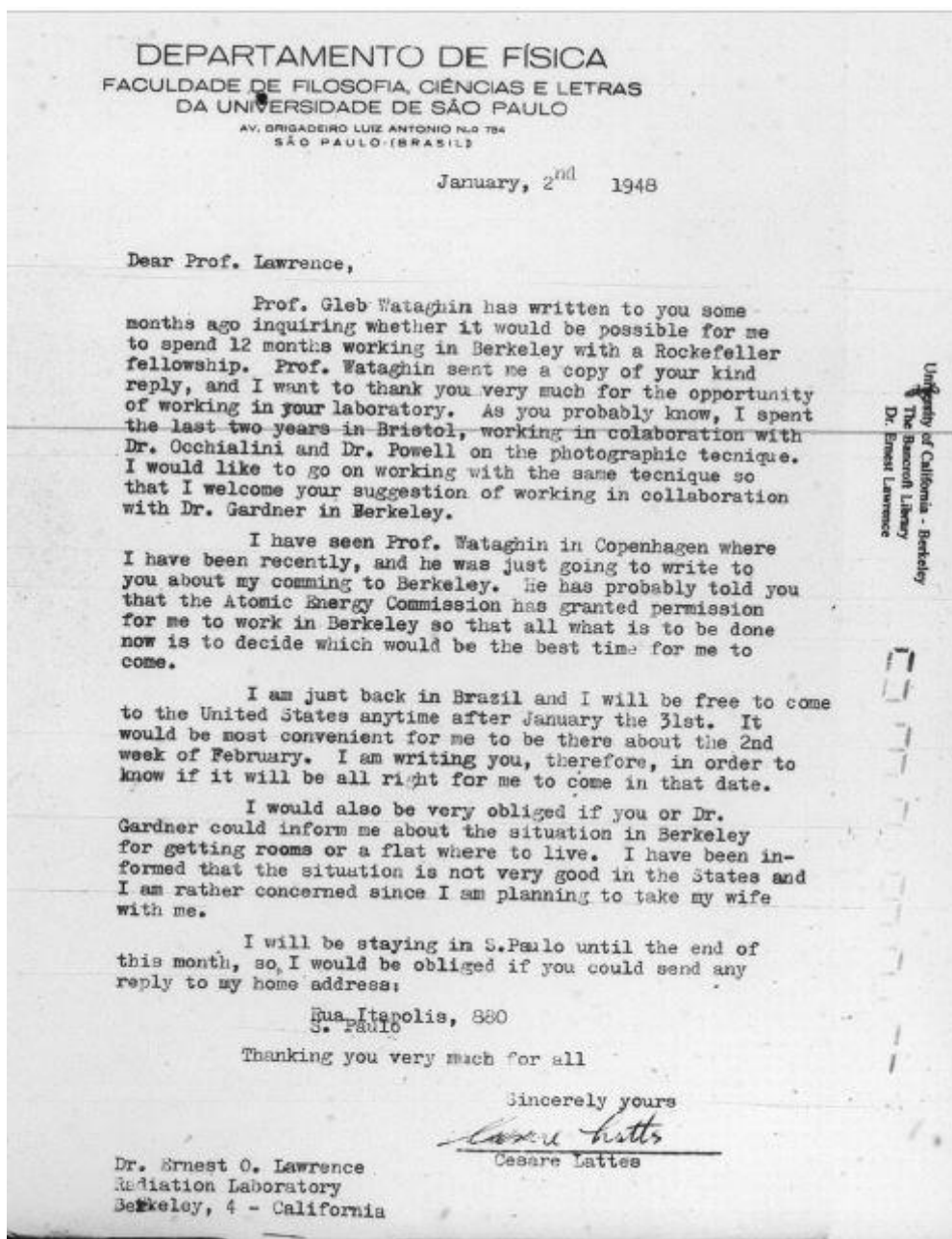
[90] C. Powell para E. Gardner (Logbooks of Meson detection experiments by Gardner Research Group, National Archives and Records Administration, San Bruno, 1947), correspondência, caixa 6, livro 10, página 188.



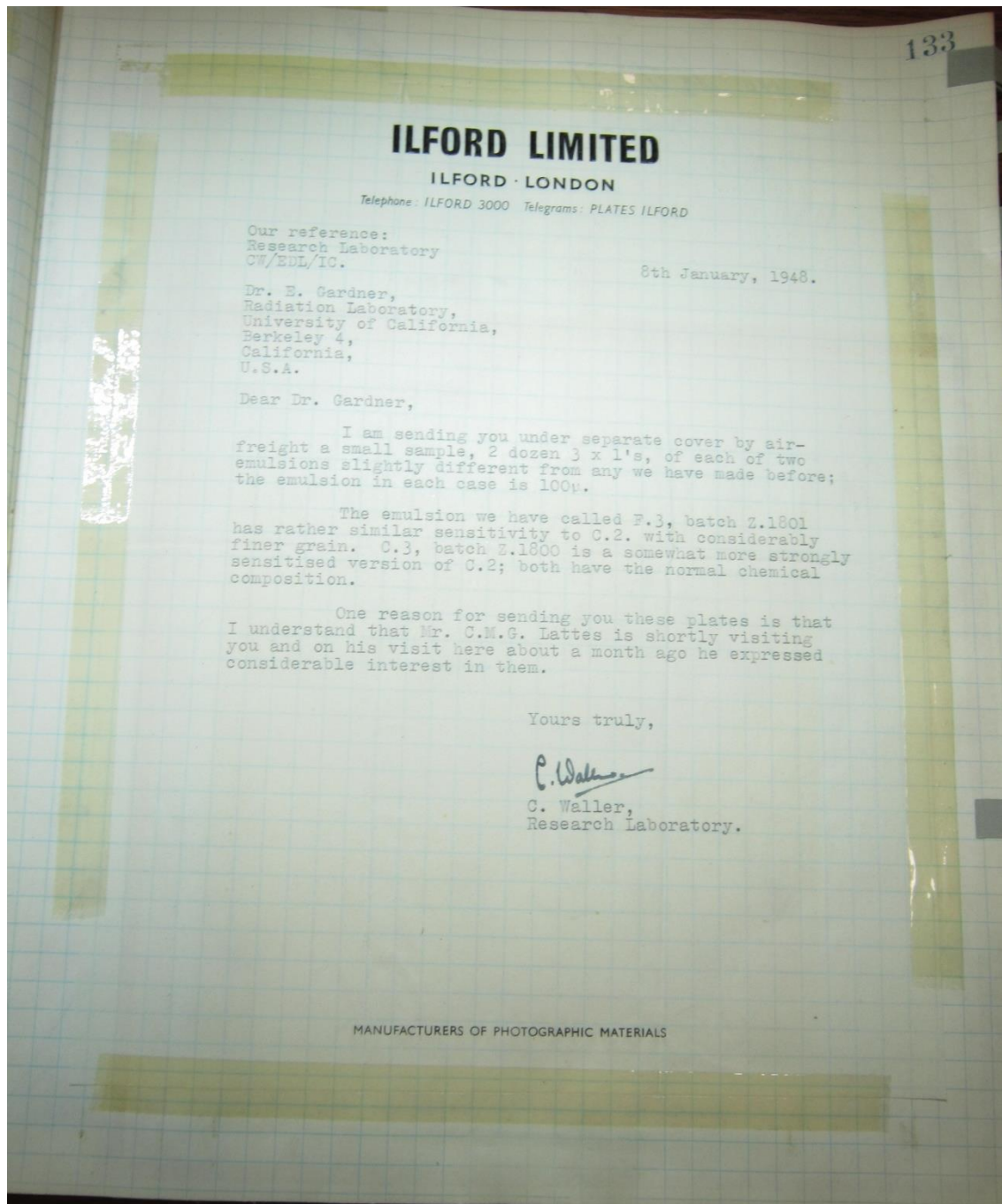
[91] E. Gardner para C. Powell (Logbooks of Meson detection experiments by Gardner Research Group, National Archives and Records Administration, San Bruno, 1947), correspondência, caixa 6, livro 10, página 215.



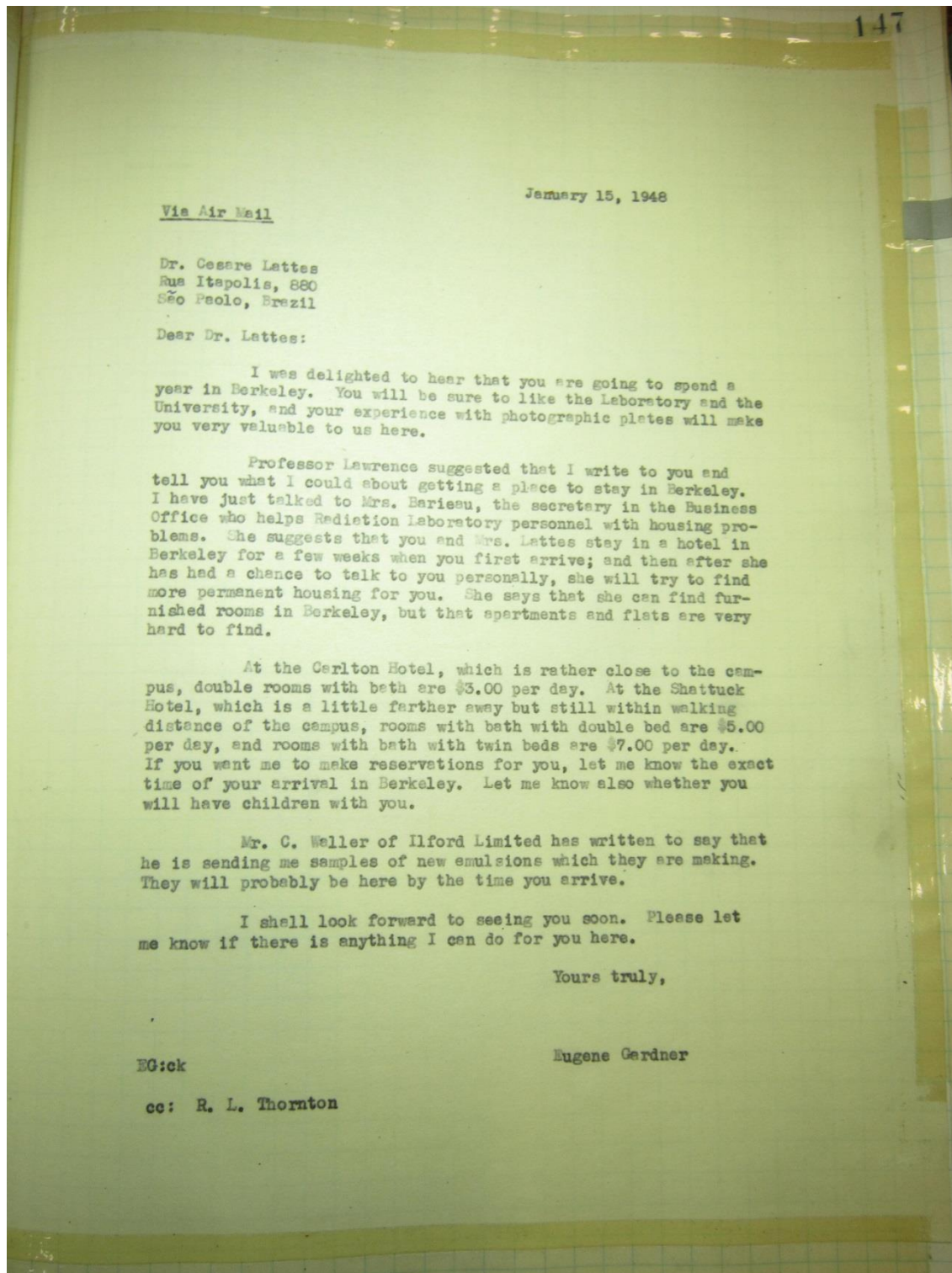
[92] C. Lattes para E. Lawrence (The Bancroft Library, University of California, Berkeley, 1948), correspondência, rolo 15, caixa 10, pasta 35.



[93] C. Waller para E. Gardner (Logbooks of Meson detection experiments by Gardner Research Group, National Archives and Records Administration, San Bruno, 1948), correspondência, caixa 5, livro 09, página 133.



[94] E. Gardner para C. Lattes (Logbooks of Meson detection experiments by Gardner Research Group, National Archives and Records Administration, San Bruno, 1948), correspondência, caixa 5, livro 09, página 147.



[95] *Special receiving record* (Logbooks of Meson detection experiments by Gardner Research Group, National Archives and Records Administration, San Bruno, 1948), caixa 5, livro 09, página 283.

S. R. R. 2

## SPECIAL RECEIVING RECORD

S. R. R. Serial No. 7700

Package Label

Shipper: Elfad Limited Shipped To: Dr Gardner Order No. \_\_\_\_\_  
 Address: \_\_\_\_\_ Shipped By: Express Dept. No. \_\_\_\_\_  
 City-State: London England F.O.B. Point: \_\_\_\_\_ Date of Order: 2/2/48  
 Other Identification: No 4 in PK

Quantity To Be Shipped	ITEM-DESCRIPTION	Property Symbol	Qt. Accepted	Actual Qt. to Shipment	Balance On Order	Remarks (Inc. Serial No.)
2	5 1" x 3" Typex C3			2.05	0	Samples. M/c.
	100 u Thick Z1800					
	Elfad Plates					
	Nuclear Research					
2	5 Nuclear Research			2.05	0	
	Typex F3-100 u Thick					
	Z1801					2/2/48

Delivered To: Dr Gardner (Person, Bldg., Room No.)  
 For Job No. 156 da 8.

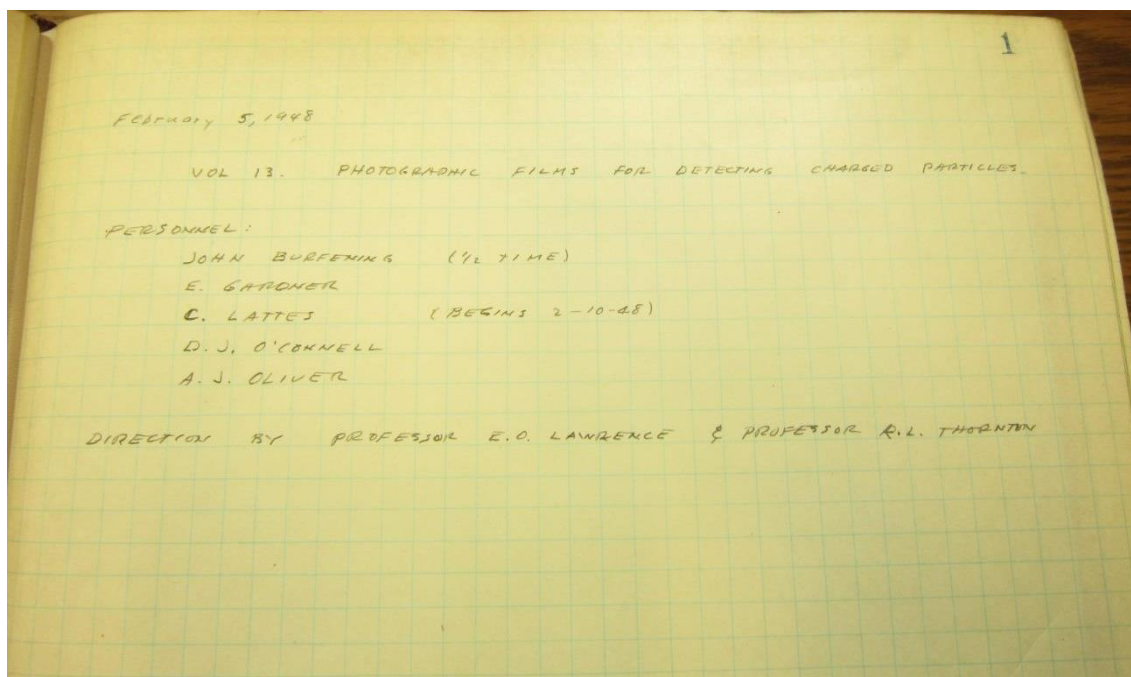
Received and Inspection Certified By: \_\_\_\_\_ Date: 2/2/48

S. R. R. Typed by

Old Town Dialform

283

[96] [Registro em 05 de fevereiro de 1948] (Logbooks of Meson detection experiments by Gardner Research Group, National Archives and Records Administration, San Bruno, 1948), caixa 7, livro 13, página 1.



[97] *Caderno de Berkeley* (Arquivo Central do Sistema de Arquivos, Universidade Estadual de Campinas, Campinas), volume 15, pp. 07, 12, 13, 14.

-1 miss on previous + ab. 1A

20th N <sup>o</sup>	B <sub>2</sub> (50 <sub>μ</sub> )	C <sub>2</sub> (50 <sub>μ</sub> )	entire 3.5 min	219	3	1	Suppos.
1	4071-4076	4089-4094	atok 4 min	D18			3 sec
2	4077-4082	4095-4100					30 sec
3	4083-4088	4101-4106					5 sec

12

Found a track very probably a muon

length - 170 small dir (x95, x10) cond  $\left\{ \begin{array}{l} 114.0 = f \\ 25.8 = x \end{array} \right.$

plate 4096

another - 210 dir cond  $\left\{ \begin{array}{l} 114.0 \\ 26.4 = x \end{array} \right.$

Negative muon K-30 Saturday - 21/2/44.

$\left\{ \begin{array}{l} x \\ 29 \end{array} \right\}; 114.0$   
(113.55)

$\rightarrow 210 \text{ dir. } \pm 5$

$\left\{ \begin{array}{l} 1 \text{ unit} = \frac{80}{75} = 1.07 \\ R = 2.6 \pm 0.1 \text{ units} \\ \text{Range} \Rightarrow 239 \pm 5 \end{array} \right.$

$\left\{ \begin{array}{l} E_{210} > 2.3 \text{ Mev interval of } 4 \left(\frac{1}{4.2}\right) \\ E_{350} > 2.8 \text{ Mev interval of } 2.2 \left(\frac{1}{5.25}\right) \end{array} \right.$

$\left\{ \begin{array}{l} x = 28.1 \\ y = 13.65 \end{array} \right.$

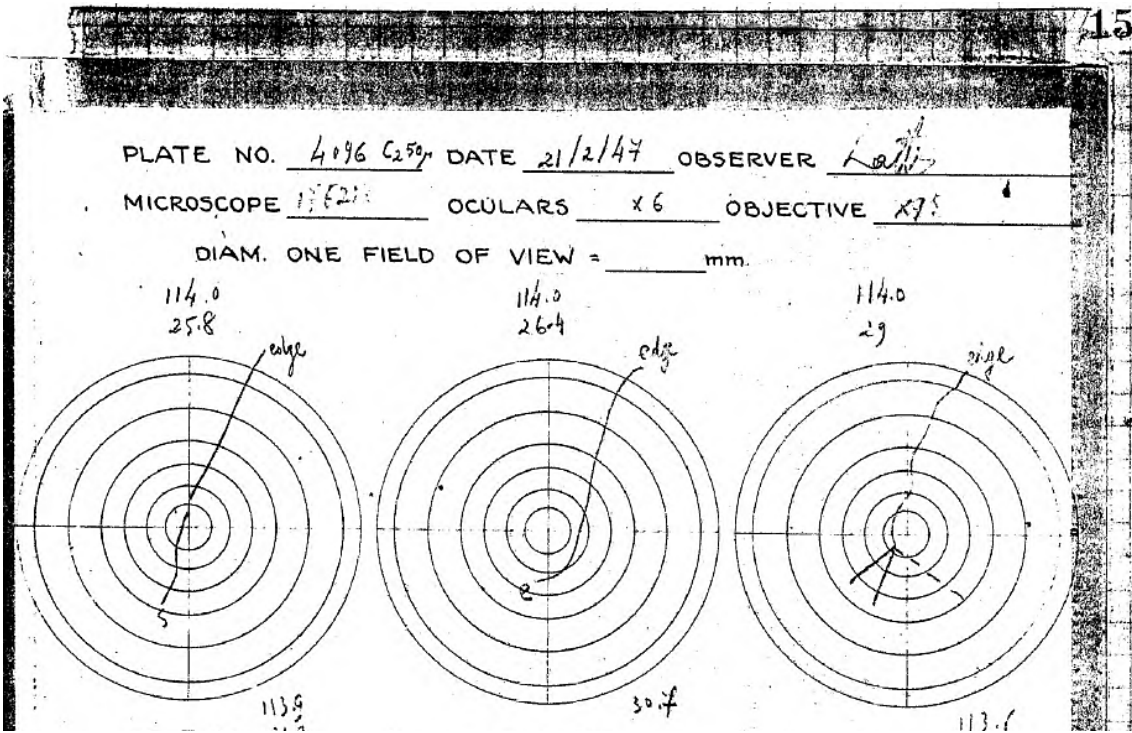
$\left\{ \begin{array}{l} 110 \\ 2120 - 2140 \end{array} \right.$

$\left\{ \begin{array}{l} 5 \\ 17 \end{array} \right.$

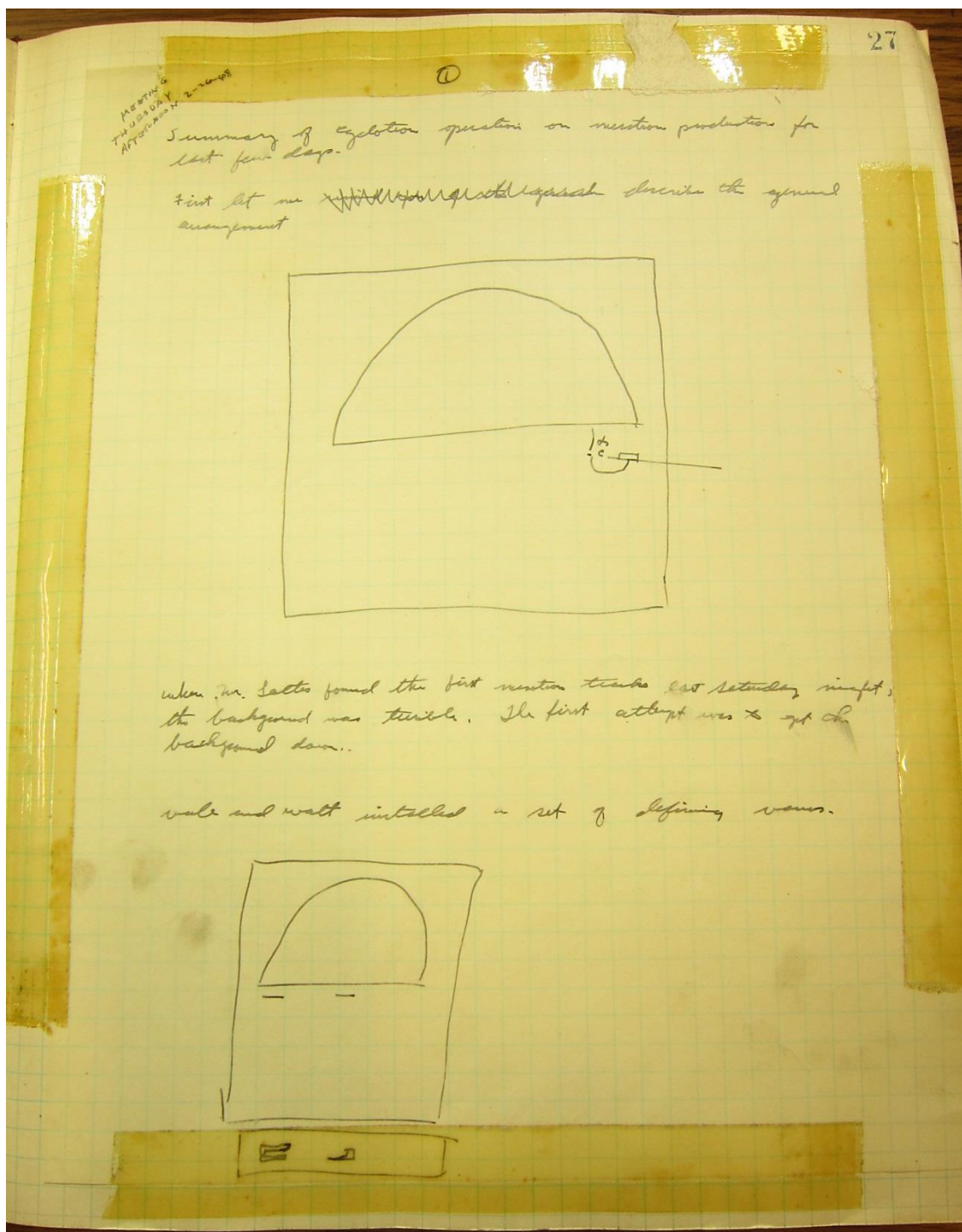
$\left\{ \begin{array}{l} 110 \\ 26.7 \\ E_6 \end{array} \right.$



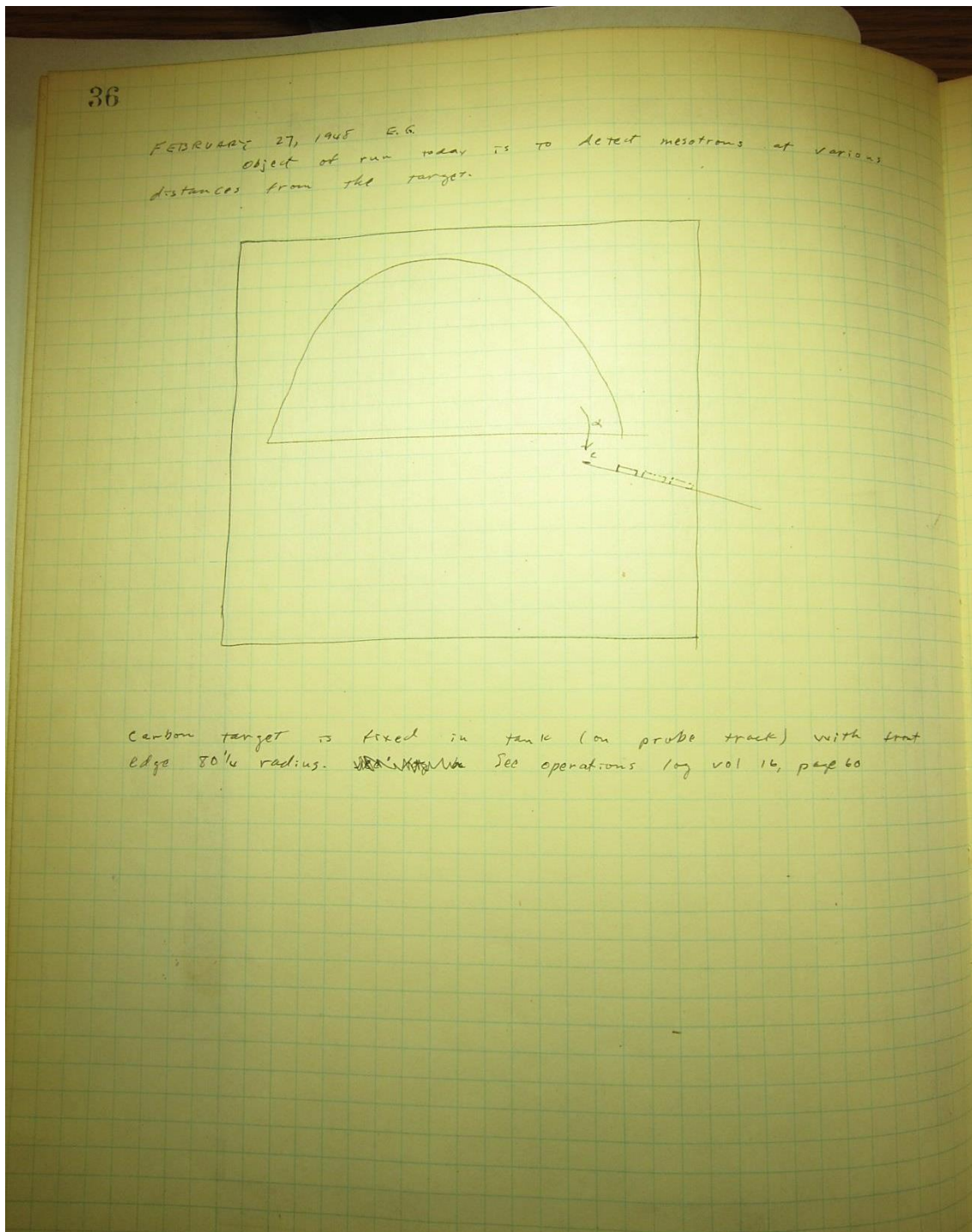
45  
 113.5 } plate 4096  
 could be protuberance (more probable)  
 30.7 micron coming from surface  
 stepping down the surface.  
 28.9 }  
 112.8 } micron very near the surface  
 goes out from slit.



[98] Meeting on Thursday (Logbooks of Meson detection experiments by Gardner Research Group, National Archives and Records Administration, San Bruno, 1948), caixa 7, livro 13, página 27.

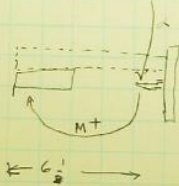


[99] [Desenho de arranjo com emulsões a várias distâncias do alvo] (Logbooks of Meson detection experiments by Gardner Research Group, National Archives and Records Administration, San Bruno, 1948), caixa 7, livro 12, página 36.



[102] *Buscas frustradas por mésons positivos* (Logbooks of Meson detection experiments by Gardner Research Group, National Archives and Records Administration, San Bruno), caixa 7, livro 12, página 27 e 31.

THIRD RUN LOOK FOR POSITIVE MESOTRON



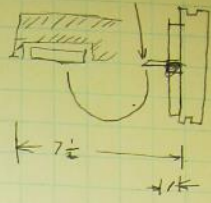
PACK NUMBER	EXPOSURE	PLATE NUMBER	TYPE OF PLATE
8	5 MIN	4449	C.2-5 $\mu$
		4450	"
9	1 MIN	4451	"
		4452	"
10	1/2 MIN	4453	"
		4454	"

Don did not find any in one day scanning.  
Very high background from ends and front.

BAD

THIRD RUN POSITIVE MESOTRONS

PACK NUMBER	EXPOSURE	PLATE NUMBERS	PLATE TYPE
19	5 Min	4471	C.2 50 $\mu$
		4472	"
20	1 Min	4473	"
		4474	"
21	1/2 Min	4487	"
		4488	"

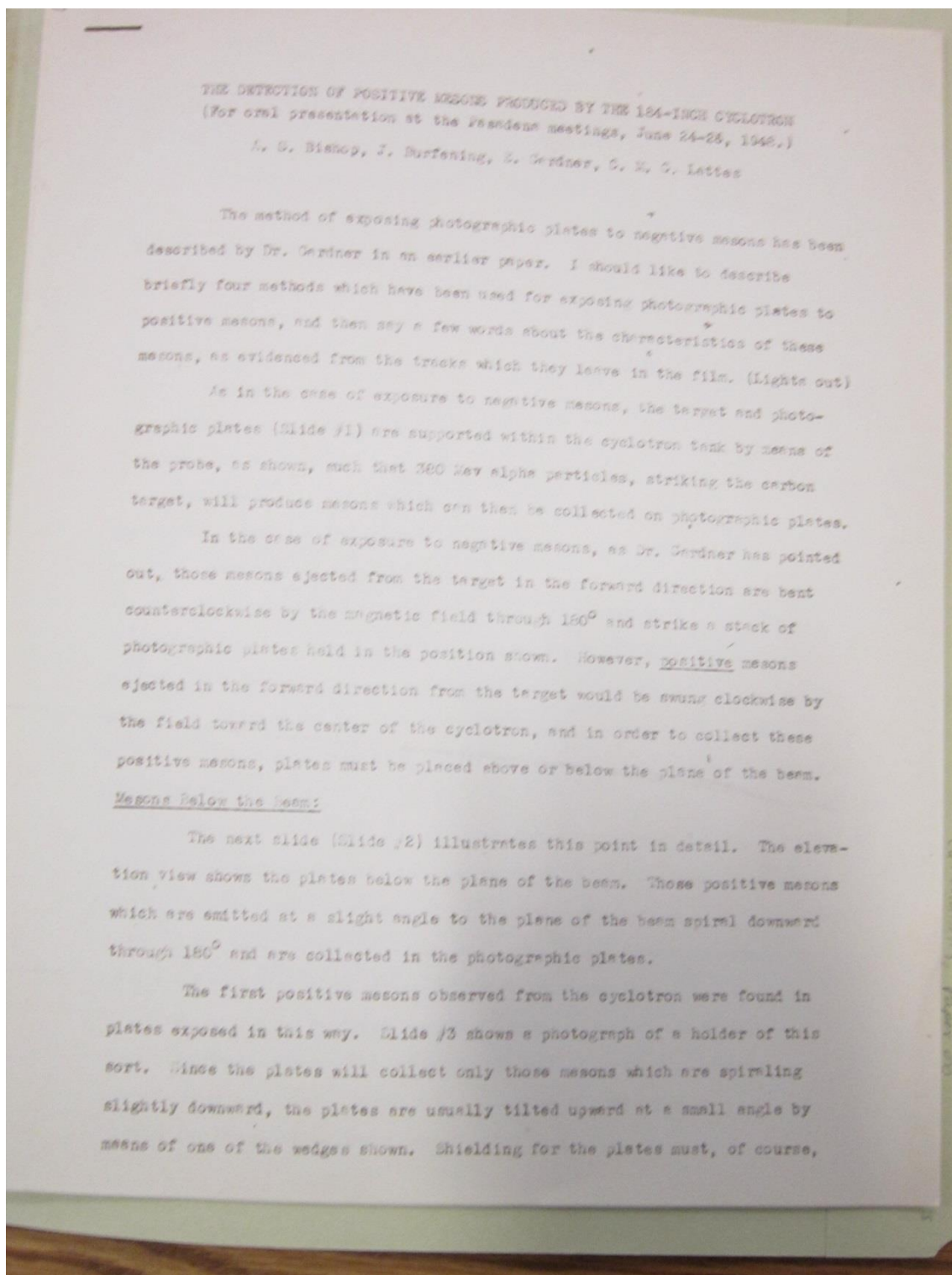


Don did not find + in one day scanning  
Background reduced somewhat.

Apparently no beam went on

4/13  
5/4  
11/3/7

[103] *The detection of positive mesons produced by the 184 cyclotron for oral presentation at the Pasadena meeting 24-28 June 1948* (Research and Development Records and Administrative Files of Eugene Gardner, National Archives and Records Administration, San Bruno, 1948), caixa 1, pasta 10.



-2-

be provided on the side from which the beam comes, in order to prevent stray fast particles in the beam from striking the plates.

Exposure to positive mesons has never been as satisfactory as exposure to negative mesons. The reason for this fact is that alpha particles and protons from the target can travel exactly the same trajectories as that traveled by the mesons. In order to prevent fast charged particles from entering the plates and causing undue background, it has been found necessary to add the copper baffle (shown in the slide) over the entrance to the photographic plates. With this baffle in position, charged particles with a large radius of curvature, i.e., low energy, can pass through the openings in the baffle. This means, of course, that low energy protons and alpha particles, as well as low energy mesons, can all enter the photographic plates. But of these particles only the mesons will penetrate appreciably into the emulsion, since the range of a meson is a number of times greater than an alpha particle or a proton with the same momentum. The alpha particles and protons which pass through this baffle are thus stopped in a few microns of emulsion while the mesons penetrate much deeper. To date the baffles have been only partially successful, however, in shutting out the high energy particles. Even with the baffles there are still so many background tracks that searching for mesons is difficult.

Mesons Ejected Backward:

Another type of holder for exposing plates to positive mesons, and that which has been most successful, is shown in Slide #4. It has been found experimentally that some mesons are emitted in the backward direction, i.e., opposite to that of the beam. Positive mesons emitted in this direction, however, can be collected very simply, in plates positioned as shown, merely by modifying the position of the heavy copper shielding to allow free passage of these mesons. Exposures made in this way allow the collection of both positive and negative mesons on the same plate; the positive mesons entering the back edge of the plate, the negative mesons entering the front edge.

-3-

270° Focusing (Slide #5)

A third, but less successful, method of exposure which has been used (and which I shall merely mention) is a variation of the method last shown. The mesons are bent through 270° and enter plates positioned as shown.

360° Focusing

The fourth method of exposure is that shown in the next slide (Slide #6). In general, each and every meson emitted from the target will follow a spiral path, due to the effect of the magnetic field, and will eventually return to a point directly above or directly below the target. Thus a small stack of photographic plates, positioned as shown, should collect the majority of mesons of both positive and negative charge, emitted over a wide range of energies and angles. Once again, to shield out the high energy alpha particles and protons, one copper shield is inserted between the target and the plates, and another in the usual position. Holders for exposing plates in this way have been built and used, but to date they have not given very satisfactory results.

Ratio of Positive to Negative Mesons

I should like to mention at this time one point of interest which is now being studied. It concerns the relative number of positive and negative mesons which are emitted from the target. To study this question plates have been exposed in a holder (Slide #7) which provides, as nearly as possible, the same geometry in the collection of positive and negative mesons. Positive mesons, spiraling in toward the center of the cyclotron, are collected in one set of plates, placed below the beam in the usual way; negative mesons, spiraling away from the cyclotron center, are collected in another set of plates which are also placed below the plane of the beam for purposes of symmetry. The wedges again are for the purpose of tilting the plates slightly upward. Here again it has been found desirable to place baffles over the entrance to the photographic plates -- on the left side in order to shut out the high energy protons and alpha particles; on the right side for purpose of symmetry in the geometry. Photographic plates have been exposed to mesons in this way and are now in the

-4-

process of being examined.

Analysis of Tracks:

Study of the plates exposed in the four ways which have been described have yielded, to date, the tracks of several hundred positive heavy mesons. Because of their charge, heavy positive mesons do not enter nuclei and make stars, but instead decay, giving rise to light mesons.

The next slide (Slide #8) shows the track of one of the first heavy positive mesons found, using the holder first shown which positions the plates under the plane of the beam. The high background is evident, particularly at the edge of the plate. Here the heavy positive meson is seen to come to rest in the emulsion. Although the track of the light meson does not show very well on this slide, it may be seen in the original plate, going off to the right and eventually leaving the emulsion. In most of the events which have been found of a heavy positive meson coming to rest, tracks of these light secondary mesons have been observed, and it is believed that in actuality all heavy positive mesons decay into light mesons. These light mesons are not always seen, however, because of unfavorable angle of emission or because the background is too high.

This meson was found in an Ilford C.2 film, 50 microns in thickness. Recently, photographic plates of 100 micron thickness have been used in order to increase the probability of finding events in which the track of the light meson remains in the emulsion throughout its entire range. Dr. Lettes and Dr. Berkes have recently succeeded in finding ten such cases, mostly in Eastman NTB plates, 100 microns thick.

The next slide shows a composite microphotograph of such an event (Slide #9). The primary meson is observed to come to rest in the upper left-hand corner, and to give rise to the secondary meson, which slows down and stops in the emulsion. No particle is observed to originate from the terminus of the light meson track. If the light meson itself decays with the emission

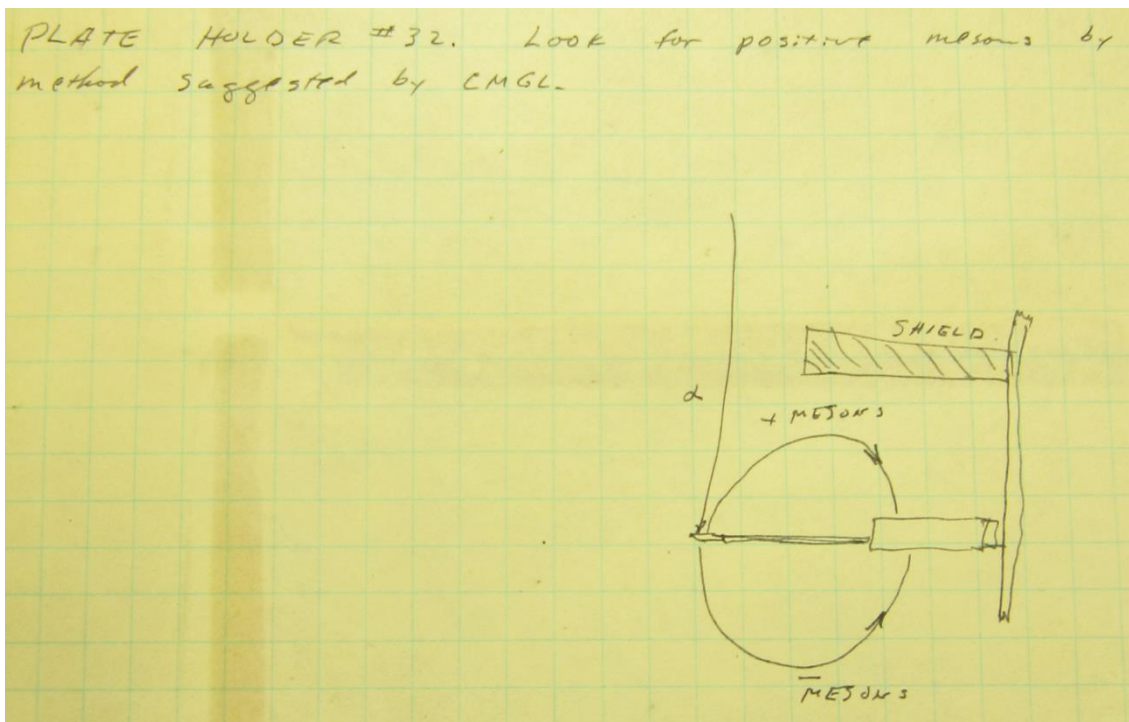
Dr. Lettes's Paper 1953



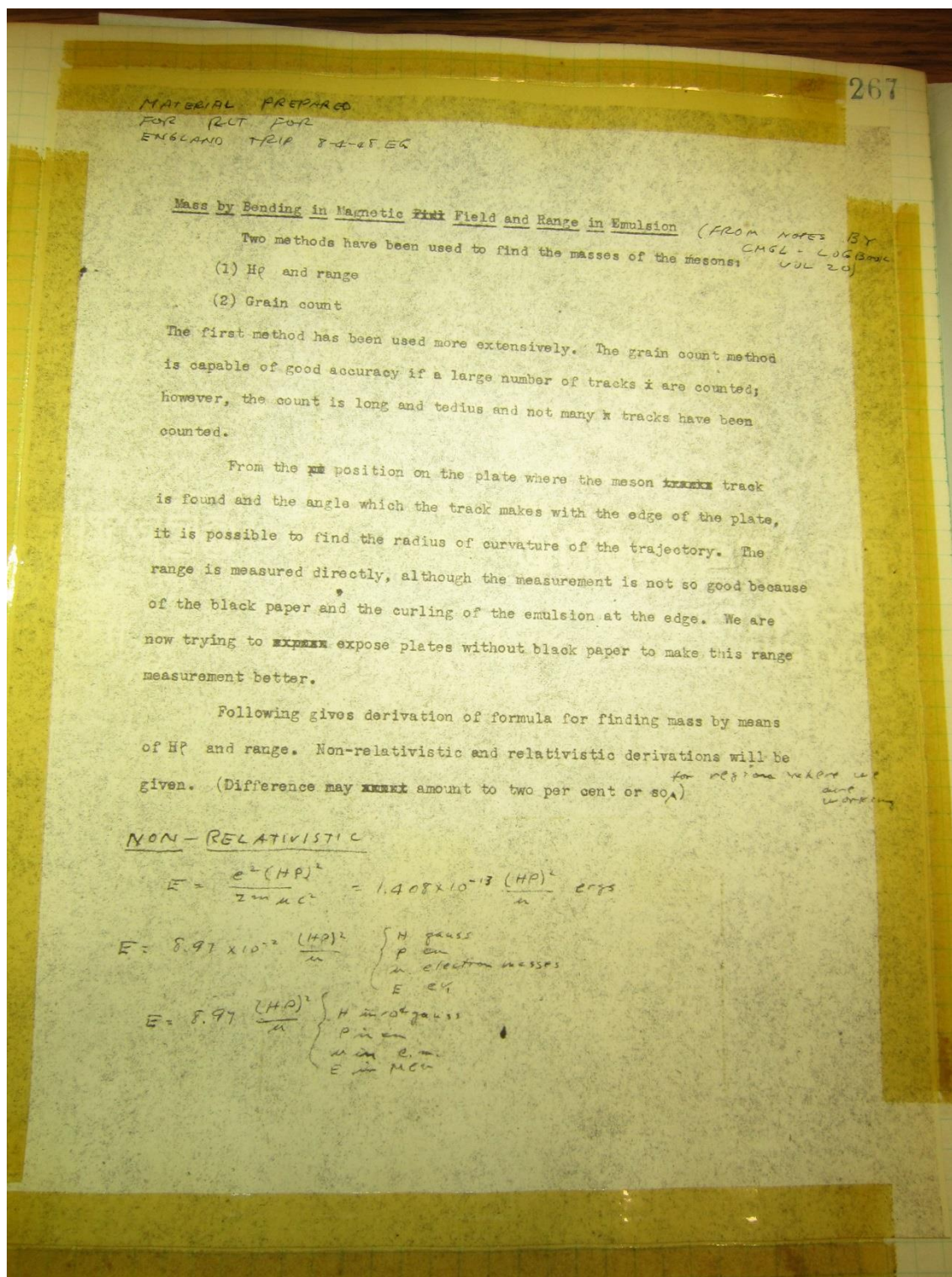
-5-

of an electron, this electron would not be observed, due to the insensitivity of the emulsion to electrons of moderate energy. The ranges of the secondary mesons, in the ten cases which have been observed, are closely distributed about 600 microns. This range agrees very well with the observations of Lattes, Occhiolini, and Powell in England on the decay of positive mesons in cosmic rays. A range of 600 microns corresponds to an initial energy of 4 Mev for the light meson, which explains the lightness of the track in the region of the decay. An enlarged photograph (Slide #10) shows that the light meson really does originate at the terminus of the heavy meson track. No track has yet been observed of any particle going off in the opposite direction to that of the light meson.  
(Lights on)

[104] Look for positive mesons suggested by CMGL (Logbooks of Meson detection experiments by Gardner Research Group, National Archives and Records Administration, San Bruno, 1948), caixa 7, livro 12, página 154.



[105] *Mass by bending in magnetic field and range in Emulsion* (Logbooks of Meson detection experiments by Gardner Research Group, National Archives and Records Administration, San Bruno), caixa 6, livro 11, página 267.



$$E = 0.305 M^{0.447} R^{0.553} \quad (\text{SEE SECTION ON RANGE-ENERGY})$$

$E$  in MeV  
 $M$  proton masses  
 $R$  microns

$$E = 1.06 M^{0.447} R^{0.553} \times 10^{-2} \quad \left\{ \begin{array}{l} E \text{ in } \mu\text{eV} \\ M \text{ in } \mu\text{eV} \end{array} \right.$$

COMBINE WITH: (FROM PAGE 1)

$$E = 8.97 \frac{(HP)^3}{M}$$

$$M = (8478) \frac{1}{R^{0.447}} \frac{(HP)^{1.382}}{R^{0.553}}$$

$$M = 171 \frac{P^{1.382}}{R^{0.382}}$$

NON-RELATIVISTIC  
 FORMULA FOR  
 MASS OF MESONS  
 IN TERMS OF  
 RADIUS OF CURVATURE  
 AND RANGE

RELATIVISTIC

$$P = \frac{E}{c} \sin \theta$$

$$\left\{ \begin{array}{l} E_T = E_C + E_0 \\ E_T^2 = P^2 c^2 + E_0^2 \end{array} \right. \quad E_0 = mc^2$$

$$E_C^2 + 2E_0 E_C + E_0^2 = P^2 c^2$$

$$E_0^2 + 2mc^2 E_C = (cHP)^2$$

$$\left\{ \begin{array}{l} \frac{(cHP)^2}{2mc^2} = \frac{E_C^2}{2mc^2} + E_C \\ E_C = \frac{1}{2} m^{1/2} R^2 \end{array} \right.$$

$$m^{1/2} R^2 = \frac{(cHP)^2 - E_C^2}{2mc^2}$$

$$m = \left\{ \frac{(cHP)^2 - E_C^2}{2A C^2 R^2} \right\}^{1/2}$$

$$m = \left\{ \frac{e^2}{2 \times c} \right\}^{0.601} (H)^{1.382} \frac{(P^2 - \frac{E^2}{c^2})^{0.671}}{R^{0.782}}$$

$$m = 171 \frac{(P^2 - 0.055 E^2)^{0.671}}{R^{0.782}}$$

$$m_A = m_c \left[ 1 - 3.8 \times 10^{-2} \left( \frac{E}{P} \right)^2 \right]$$

↑  
mass  
in relativistic  
calculator

↑  
mass on  
non-relativistic  
calculator

↑  
E in Mev  
m in cm  
P in cm

RELATIVISTIC  
FORMULA  
FOR MASS  
OF MESONS  
IN TERMS  
OF RADII  
OF CURVATURE  
AND RANGE

- END -

[106] Range energy (Logbooks of Meson detection experiments by Gardner Research Group, National Archives and Records Administration, San Bruno), caixa 6, livro 11, página 270.

70

Range-Energy (FROM NOTES BY ENGL - LOG BOOK VOL 30)

Range-energy curve from Bristol data. Reactions of light elements.  
 \*Lars. Fowler, Chem. Phys. Ser. Vol 57, Sept. 47  
 Get good fit to experimental range-energy curve by following formula:

$$E = 0.305 M^{0.447} R^{0.553}$$

$E$  = energy, Mev  
 $M$  = mass in proton masses  
 $R$  = range, microns

EXAMPLE: Find energy of protons of range 56.4  $\mu$   
 $E = 0.305 \cdot 1^{0.447} (56.4)^{0.553}$   
 $= 0.305 \cdot 3.3 = 1.0 \text{ Mev.}$

The derivation of this formula is as follows: The Bristol data on light element reactions give 13 points, all at energies less than 13 Mev. (protons)

~~The points are used to find NZ and I in following formula~~ (For use with mesons the proton energies of interest are considerably higher than this) The points are used to find NZ and I in following formula (Good at low energies only)

$$-\frac{dE}{dR} = \frac{4\pi e^4 z^2}{m v^2} NZ \left( \frac{m v^2}{I} \right)$$

Find: NZ (experimental) =  $0.94 \times 10^{24}$  per  $\text{cm}^3$   
 (theoretical) =  $1.09 \times 10^{24}$  per  $\text{cm}^3$

(The theoretical value is found by taking all of the atoms in the emulsion, multiplying by Z. The fact that the theoretical value is larger suggests that not all of the ~~min~~ electrons are effective.)

Also find:  $I = 231$  e.v.

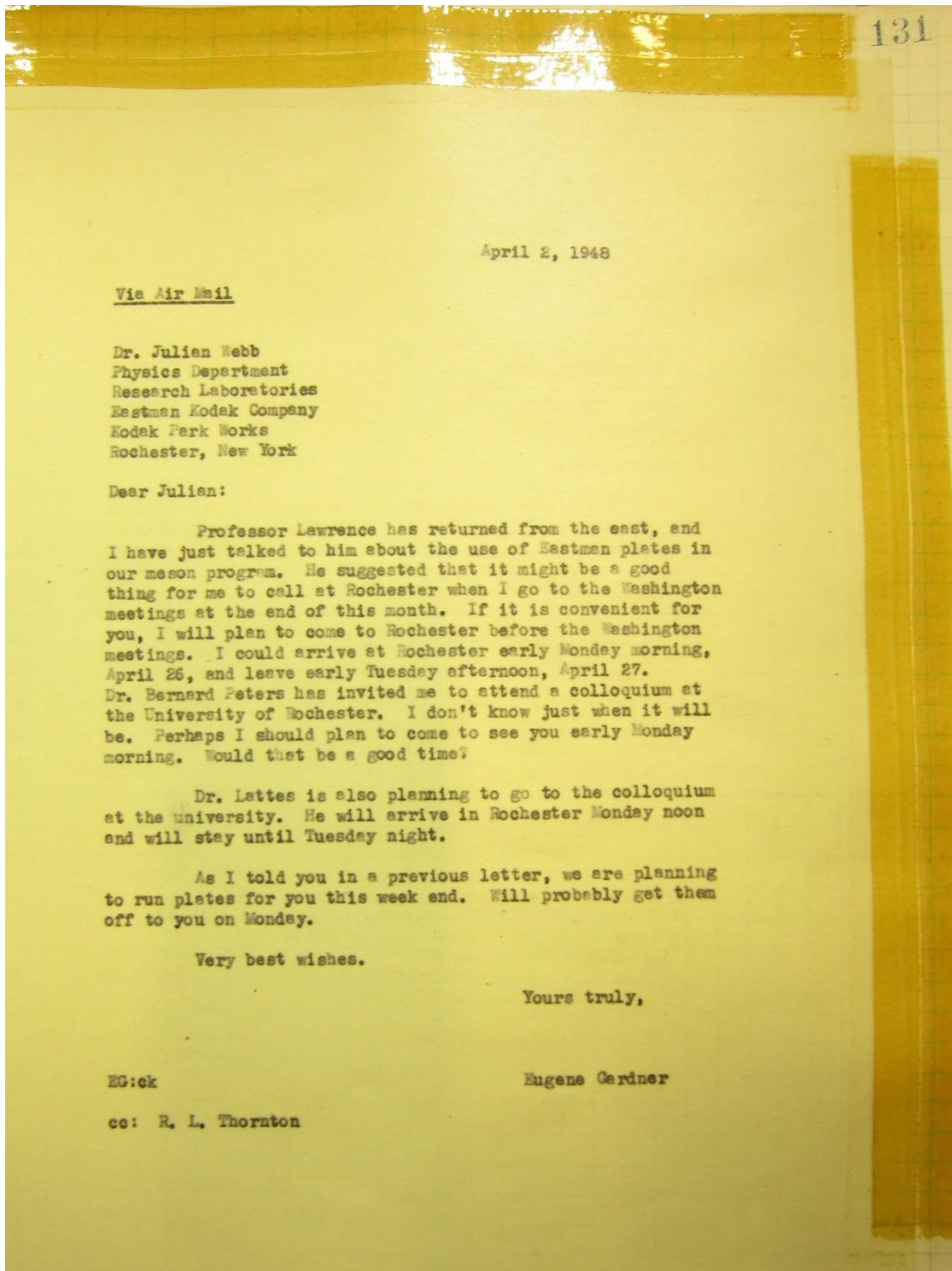
Now use these values of NZ and I in formula which is good at higher energies.

$$-E \frac{dE}{dR} = \frac{4\pi e^4 z^2 M}{m} NZ \left\{ \log \left( \frac{dEm}{Im} \right) - \log (1-\beta) H^2 \right\}$$

Solve for R by numerical integration. The result is a table of values for range as a function of energy. This table of values is now fitted by best equation for E as function of mass and range. ~~Result~~ Result is equation top page 1. Equation is estimated to be correct to within  $\pm 2$  per cent. It is found to check the range energy curve within 0.4 per cent.

One point from the linear accelerator (32 Mev protons) was measured (gave 4.55 mm). ~~Fits~~ Fits equation well within the 2 per cent estimated error of the equation.

[107] E. Gardner para J. Webb (Logbooks of Meson detection experiments by Gardner Research Group, National Archives and Records Administration, San Bruno, 1948), correspondência, caixa 7, livro 13, página 131.



April 2, 1948

Via Air Mail

Dr. Julian Webb  
Physics Department  
Research Laboratories  
Eastman Kodak Company  
Kodak Park Works  
Rochester, New York

Dear Julian:

Professor Lawrence has returned from the east, and I have just talked to him about the use of Eastman plates in our meson program. He suggested that it might be a good thing for me to call at Rochester when I go to the Washington meetings at the end of this month. If it is convenient for you, I will plan to come to Rochester before the Washington meetings. I could arrive at Rochester early Monday morning, April 26, and leave early Tuesday afternoon, April 27. Dr. Bernard Peters has invited me to attend a colloquium at the University of Rochester. I don't know just when it will be. Perhaps I should plan to come to see you early Monday morning. Would that be a good time?

Dr. Lettes is also planning to go to the colloquium at the university. He will arrive in Rochester Monday noon and will stay until Tuesday night.

As I told you in a previous letter, we are planning to run plates for you this week end. Will probably get them off to you on Monday.

Very best wishes.

Yours truly,

EG:ck

Eugene Gardner

cc: R. L. Thornton



[108] *Conference on nuclear particle plates* (The Bancroft Library, University of California, Berkeley, 1948), rolo 13, caixa 06, pasta 25.

CONFERENCE ON NUCLEAR PARTICLE PLATES

April 26, 1948

Present: Dr. Eugene Gardner, University of California; Mr. E. B. Davies of Kodak Limited; Dr. C. E. K. Mees, Dr. C. J. Staud, Dr. J. A. Leermakers, Dr. W. Vanselow, Dr. W. West, Dr. E. H. Carroll, Mr. W. F. Swann, and Dr. J. H. Webb of Eastman Kodak Company

Dr. Staud opened the meeting by asking Dr. Gardner to tell us a little about the recent experiments on the observation of laboratory produced mesons at the University of California.

Dr. Gardner made a few remarks about the experiments in which the mesons were observed and then discussed mainly the use of the photographic plates in their experiments. Dr. Gardner stated that the mesons had been observed using Ilford G2 plates and that these plates were exposed and processed under the direction of Dr. C. M. G. Lattes who was well versed in the most up-to-date techniques for handling the Ilford plates to get the most from them. Dr. Lattes recently returned from Bristol University where much work has been done on the observation of cosmic ray mesons. In Bristol, both heavy (313 Me) and light (200 Me) mesons were studied by means of the photographic plate, and Ilford plates had shown ample sensitivity to register considerable path lengths for these particles. Dr. Gardner pointed out that the development recommended by Dr. Lattes was considerably longer than the development that had previously been given Ilford plates by the University of California Laboratory and that this probably accounted for the fact that they had not observed mesons on the Ilford plates previous to Dr. Lattes' visit.

Dr. Gardner pointed out that Eastman plates had been used to look for mesons but without success and that they had attributed this failure to lack of sensitivity of the Eastman plates or to improper processing conditions. However, it was pointed out that the trials made with Eastman plates had not been extensive and furthermore that the plates that had been used were several months old and certainly not recent NTB plates that are known to be capable of registering mesons (3-5 Mev) and electron tracks (50 KeV). Dr. Gardner pointed out that it had been difficult to obtain Eastman plates without considerable delay and that it had been necessary for them to buy fairly large quantities and store them for future use.

Dr. Gardner was very emphatic in pointing out that their experience with the Eastman nuclear track plates had shown them to be variable in their response to high-speed particles. He said that some batches of the Eastman plates had been very insensitive. Dr. Gardner said he thought that the manufacturer of

nuclear track plates should be very frank in telling scientists the true qualities of these plates and that if they are inconstant in their properties, the scientists should be told so. Dr. Gardner pointed out that scientists were prone to take a plate and use it assuming that it were absolutely constant in its properties. Therefore, he thought it was up to the manufacturer to point out to the scientist the variations to be expected so that he would not be misled and thereby publish incorrect conclusions. Everyone agreed with Dr. Gardner that frank information should be given on the plates.

Dr. Spence pointed out that the NTB plates had been somewhat variable in their properties. However, these plates have always been sold on an experimental basis and sold under the label of an experimental product. He said that the experimentation on these plates had been the cause of some changes in the plate, but generally speaking, the sensitivity had been on the upgrade all the way through.

Dr. Gardner pointed out that Eastman plates in some ways were superior to Ilford plates, e.g., in having lower background.

Dr. Gardner passed around pictures of meson tracks obtained on Ilford plates at the University of California, and Dr. Spence passed around pictures of meson tracks obtained on the latest NTB Eastman plates (exposed at California and processed here at Rochester). Dr. Spence also showed some electron tracks of about 20 grain lengths obtained on our recent NTB plates. Everyone showed great interest in the meson tracks obtained on the Ilford plates in California, and Dr. Gardner was pleased with the meson tracks of 400  $\mu$  length shown on the recent NTB emulsion.

Dr. Spence raised the question of devising good tests for our NTB plates. This test should be easy to perform, repeatable, and adequate for determining the limit of sensitivity for low ionizing particles. All of these criteria seem to be met by a test using electron tracks. Photoelectrons obtained by irradiation with known voltage x-rays can be used for exposing the plates. The length of electron tracks obtained as well as the density of grains in the tracks will serve as a testing procedure that can be carried out here. Mr. Davies suggested that a test might be devised whereby direct density measurements of the x-ray exposures can be used instead of looking at tracks. The exposure in the case of x-rays being mostly due to electrons, density measurements might be sufficient for measuring sensitivity. The general feeling was, however, that in addition to a density measurement, it would also be necessary to observe some tracks.

Considerable discussion was devoted to the subject of standardization of the NTB plate. At present this plate is sold as an experimental plate in various thicknesses and sizes. The consensus of opinion seemed to be that we should standardize this plate as a fixed commercial item.

Size	1" x 3"
Thickness	50 $\mu$
Sensitivity	To show electron tracks to 50 KV.

Dr. Spence pointed out that 12 good batches of these plates had already been made and therefore we should have no trouble in reproducing them now. It was further agreed that we should sell these plates in any size up to 8" x 10" and

in thicknesses up to 100  $\mu$  on special order. However, plates on special order would not be delivered as promptly as the standard plates. It was suggested by Mr. Swann that all of this information should be included in the pamphlet on these plates that is now being prepared.

The matter of quantity production of the NTB plates was discussed at some length. Dr. Gardner stated that the Radiation Laboratory of California could probably get along with 5000-1" x 3" plates per year. He pointed out that in the near future there would be a number of other accelerators coming into use over the country and that each one would probably use as many plates as the University of California. It appeared, after some discussion, that a good production figure to aim at would be 100,000-1" x 3"--50  $\mu$  thickness-- plates per year. It was further agreed that we should plan to build up a stock of these plates so that orders can be filled promptly.

Mr. Swann asked what the shelf life of the plates is. Dr. Spence stated that the plates would keep 6 months at 70°F. However, the plates should be refrigerated for better keeping, and it was suggested that we advise the users to keep the plates in deep-freeze units. This type storage would not only preserve the plates against fog but to some extent against fog from cosmic rays and radioactive contaminants in the glass plates.

There was some discussion among the Kodak people concerning the facilities for manufacturing 100,000-1" x 3" plates per year. It was pointed out that if the spectroscopic plates are moved from Building 3, then there will be additional space available for the nuclear particle plate work. Dr. Staud said that possibly production could be increased by putting three shifts to work on the production of plates. Dr. Spence was asked to look into the possibilities for doing this.

Dr. Leermakers asked Dr. Gardner if he could tell us about the uniformity of Ilford plates in connection with their use on the Cyclotron.

Dr. Gardner said that they had recently received some bad batches of Ilford plates--bad fog without exposure and variations in sensitivity from batch to batch. Dr. Gardner was of the opinion that Ilford makes their emulsions in very large batches. Variations within a single emulsion batch are very small.

It was suggested that we make NTB emulsions in large batches by blending several small batches of emulsion to make one large batch.

Dr. Spence said that this procedure was being followed at present.

Dr. Gardner remarked that in addition to the regular production of standard plates, they would like special experimental products from time to time--he referred to the plates that Dr. E. O. Lawrence had requested earlier impregnated with various elements. Dr. Gardner said that at the present time, e.g., they wanted plates with heavy metals distributed in small discrete specks in the emulsion, e.g., gold, lead, etc., that would act as targets in the emulsion. Such small specks of metal hit by mesons would be the centers of stars and would enable the experimenter to see the type of star originating from particular types of atoms.

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The Bancroft Library  
Dr. Ernest Lawrence

It was pointed out to Dr. Gardner by Dr. Mees that we can probably incorporate metal in an emulsion by first dispersing the metal in an oil emulsion and then putting this into the photographic emulsion. The oil acts as a protective agent around the metal. It was suggested that Dr. Carroll be consulted on this job.

Dr. Gardner said that we should talk directly to Dr. C. W. G. Lattes about this problem, since he was the one mainly interested and he is the one who will use the plates.

The question of development of our NTB plates was brought up for discussion by Dr. Leermakers.

Dr. Gardner said that they had missed observing mesons on the Ilford plates earlier by not using the recommended long-time development. He added that he did not know what would be the effect of prolonged development on Eastman NTB plates.

It was pointed out by Dr. Spence that the good meson tracks on Eastman NTB plates shown earlier at this meeting were obtained with 2'-development in D-8 developer. He added that 5'-development would show only very slight improvement.

Dr. Staud said that Dr. A. Newton would spend 2 weeks at the Research Laboratory (May 3-17) studying development condition of the nuclear track plates and that he would convey all information thus gained back to Berkeley.

Dr. Leermakers pointed out further that we are getting a woman scientist from the University of Wisconsin for the coming summer to study the processing conditions of plates. She will be working with Dr. W. Vanselow.

It was suggested by Dr. Mees that further work might profitably be done on the microscopy of images of tracks. In this connection, he referred to the work of Burch at Oxford who had recently published work on the reflecting microscope, and it was suggested that we look into this work.

Dr. Gardner remarked that the Eastman Laboratory had made some nice emulsions for him with varying AgBr content for the study of star formation. This work was carried out for the purpose of ascertaining what elements of the plate are mainly responsible for star formation. He expressed thanks to the Eastman Laboratory for preparing these plates.

Dr. Spence mentioned that different impregnating compounds can now be put in our plates including: Boron, Lithium, Beryllium, and Bismuth. Where we incorporate these elements in the emulsion, we specify the processing conditions. However, it was pointed out that when the customer does his own impregnating of plates, it should be up to the customer to determine his own processing condition.

Dr. Staud summed up the main points of the meeting as follows:

#### Production

1. We have so far made 12 good batches of the high sensitivity NTB plates that record good meson tracks and 20 grain electron tracks. We will go

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Dr. Ernest Lawrence

ahead and try to standardize on this plate. We will build up a stock file of these plates with the idea of filling orders in future. We won't make any statement right now about delivery dates.

2. We shall aim to produce large quantities of this emulsion--10,000 dozen (1 x 3) per year.

3. We shall standardize on 1" x 3" plates, 50  $\mu$  thick.

4. We shall supply any size plates up to 8" x 10" and thickness up to 200  $\mu$  on special order. We shall state in our literature that these special plates will be slower in delivery.

#### Tests

1. We shall try to work out a standard electron track exposure for testing our plates. We shall look into the matter of the x-ray-density type test suggested by Mr. Davies to see if this is a good test. It appears that it might be, since most of the x-ray exposure is due to electrons.

2. We now have 50 dozen of the Harrow high sensitivity proton plates for distribution to scientists. These will be distributed 1 dozen to each customer free. At the same time, we shall send one dozen NTB plates and tell the customer to try both and let us know which he likes best. In this way, we shall get information on both the Harrow and Rochester plates for comparison.

#### Processing

1. Dr. Newton will spend two weeks in Rochester working on processing NTB plates, and he will take the results of this study back to Berkeley.

2. We are getting a woman scientist to work with Dr. Vanselow during the coming summer, and she will spend some time on development.

#### Observing Tracks

1. We shall look into the matter of the Barch reflecting microscope to see if this has any advantages for observing tracks.

#### Research Work

1. We shall do follow-up work on the variable AgBr content emulsions already sent to Dr. Gardner. Dr. Gardner is still working on the plates we sent him earlier.

2. We shall discuss with Lattes and do work on the problem of incorporation specks of heavy metals in the emulsion.

3. Dr. Spence will continue to experiment in an attempt to push sensitivity of the nuclear track plates up to the ultimate limit. We should like to send experimental plates to Dr. Gardner for tests--possibly meson and electron exposures. This must be worked out between Gardner, Spence, and Webb.

JHWebb:msm

*J.H. Webb*

[109] Cy para E. Lawrence (The Bancroft Library, University of California, Berkeley, 1948), correspondência, rolo 13, caixa 06, pasta 25.

EASTMAN KODAK COMPANY  
ROCHESTER 4, N. Y.

May 24, 1948

Dr. E. O. Lawrence  
Radiation Laboratory  
University of California  
Berkeley 4, California

Dear Ernest,

*Arrived  
5-29-48*

Mr. E. R. Davies, who is director of the Harrow Research Laboratory of Kodak Limited, has been in Rochester for some weeks and, before returning to England, is making a rather hurried trip to the west coast. He is expecting to arrive in Los Angeles this morning, May 24, and according to his present plan will reach San Francisco Saturday evening, May 29, and expects to leave Sunday evening, May 30. Mr. Davies would like very much to see your cyclotron, if arrangements could be made, and if it would be any more convenient for you, I think that he could cut his stay in Hollywood by one day and arrive in San Francisco on Friday evening, May 28.

I do not know whether the laboratory would be open, or whether there would be any regulations concerning the admission of aliens from a security standpoint. It occurred to me that if the admission aspect was clear, possibly Amos Newton might be able to show Mr. Davies about for an hour or two on the hill. I am sure that Mr. Davies would appreciate an opportunity to say hello to you during his brief stay in San Francisco, if this can be arranged conveniently.

We enjoyed having Amos Newton here, and appreciated the visit of Dr. Gardner and Dr. Lattes. They are both fine men, and I know you must feel fortunate to have them on your staff.

It appears to us that the NTB plates are now generally useful where they apparently had some shortcomings earlier, and we are planning to produce plates with a given formula rather than carry our experimentation into the cyclotrons. However, we hope and expect to be able to send experimental plates to Berkeley as part of our joint cooperative program. Dr. Lattes made some suggestions for types of plates to be used in specific problems in which he is interested, and those of us concerned here you may be sure will do our best to provide Dr. Lattes the best plates we can make for his investigational work.

Dr. E. O. Lawrence -2-

May 24, 1948

We have issued a small mimeographed pamphlet to various users of nuclear track plates, which we hope will get our information into a concise form and make it generally available. I do not know whether Amos Newton took back a copy of it with him when he left, but I am attaching one to this letter.

We are still trying to get a lawn started at 260 Kilbourn Road but, at the present rate, the color of the yard is not changing rapidly from brown to green.

Both May and I recall with pleasure your last visit here, even though it called for dinner in 12.45 minutes. We are looking forward to your next visit, and hope that on one of these occasions we may be able to work out a better schedule than the early evening train which you take to Schenectady.

With best regards to Molly and yourself,

Sincerely yours,



Research Laboratories

CJstaud/LM

University of California - Berkeley  
The Bancroft Library  
Dr. Ernest Lawrence

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